



LISA Pathfinder: Mission Status



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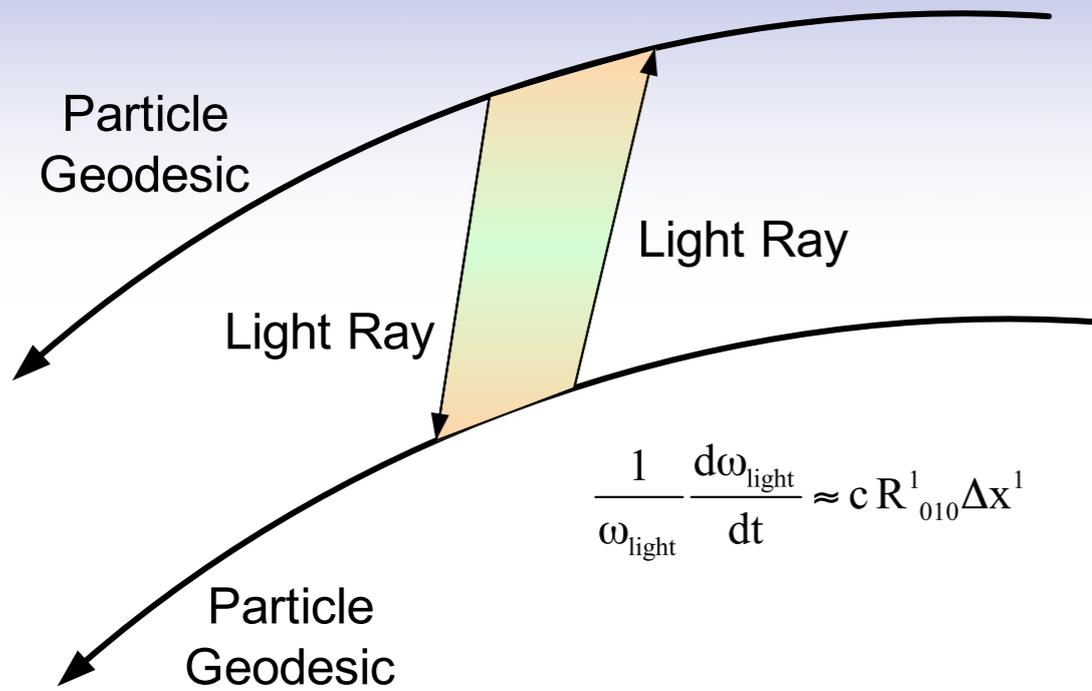
Introduction

-  LISA Pathfinder (LPF) is a technology validation mission for NGO¹
 - LPF was approved by ESA to demonstrate the concept of low frequency gravitational wave detection in space
-  LISA Pathfinder will test in flight:
 - Inertial sensors (a.k.a. Gravitational Reference Sensor)
 - Interferometry between free floating test masses
 - Drag Free and Attitude Control System (DFACS)
 - Micro-Newton propulsion technology
-  The basic idea of LISA Pathfinder is to squeeze one arm of the NGO constellation from 1 million km to a few tens of cm!
 - Fully tests NGO local interferometry



¹ When approved by the Science Programme Committee, LISA Pathfinder was a dedicated technology validation mission for the LISA mission. In the mean time, LISA has been reformulated and renamed NGO

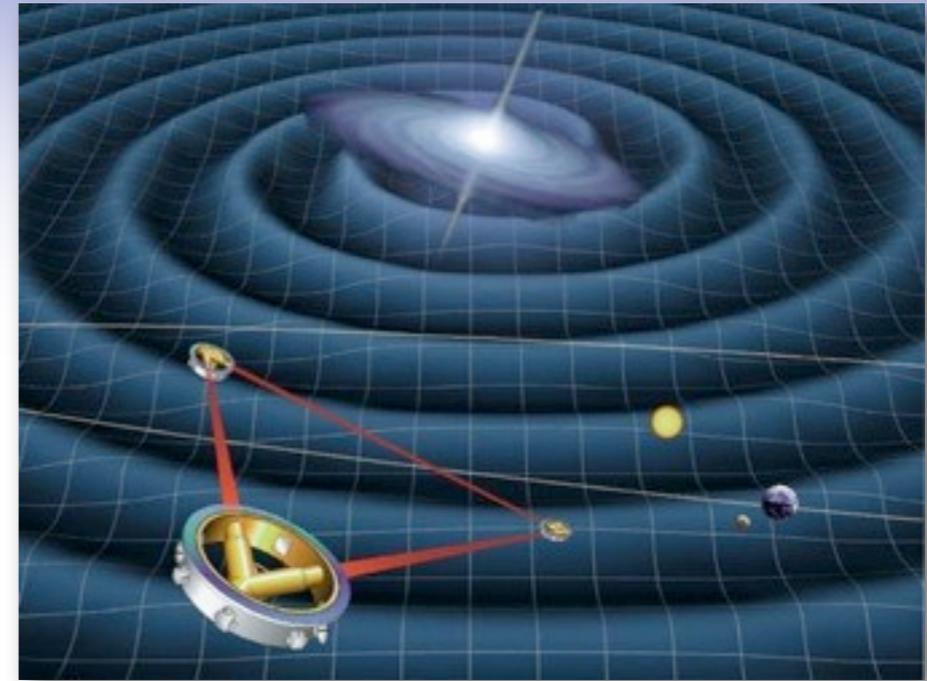
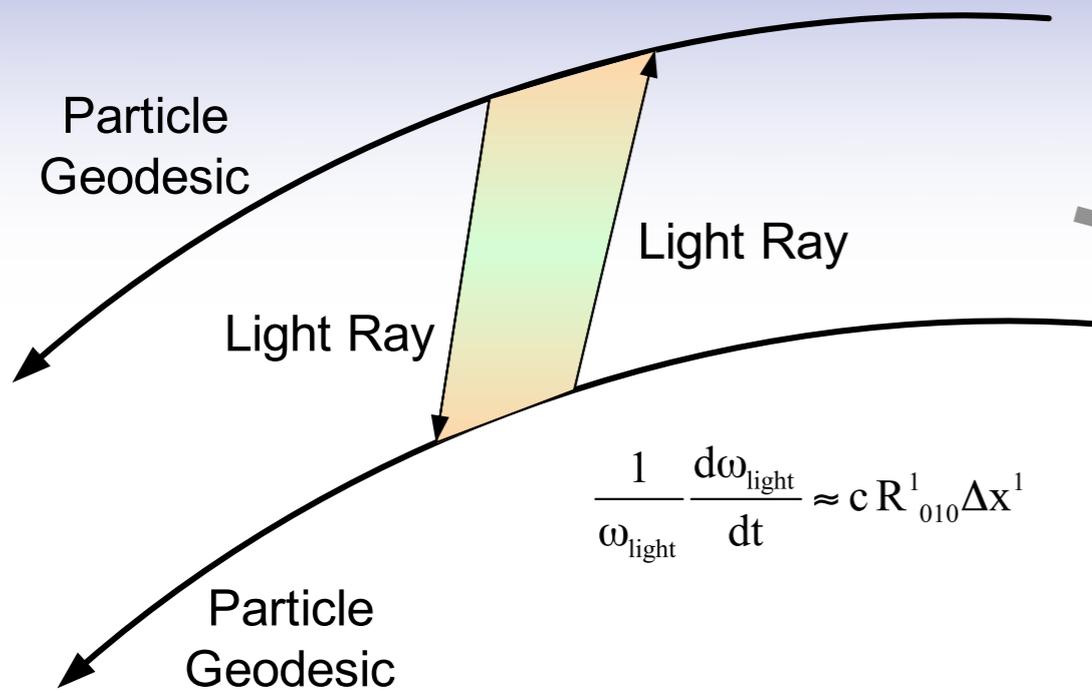
Mission Concept



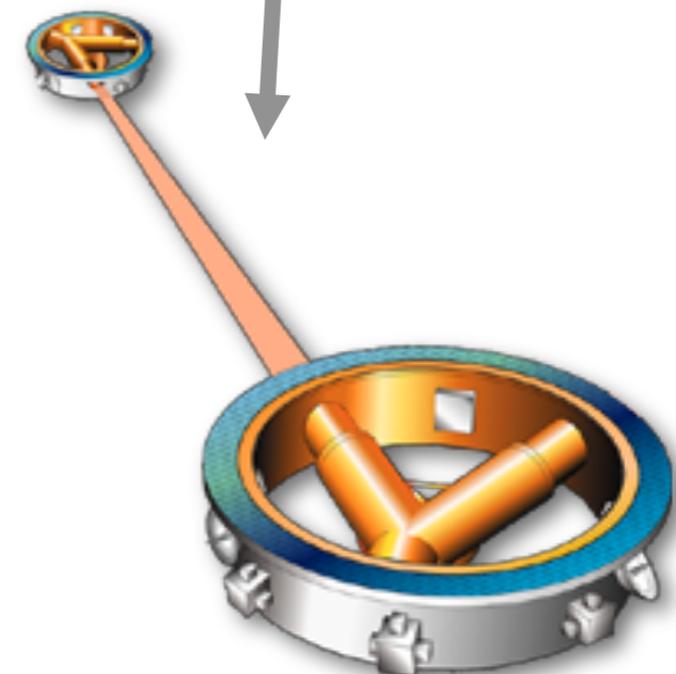
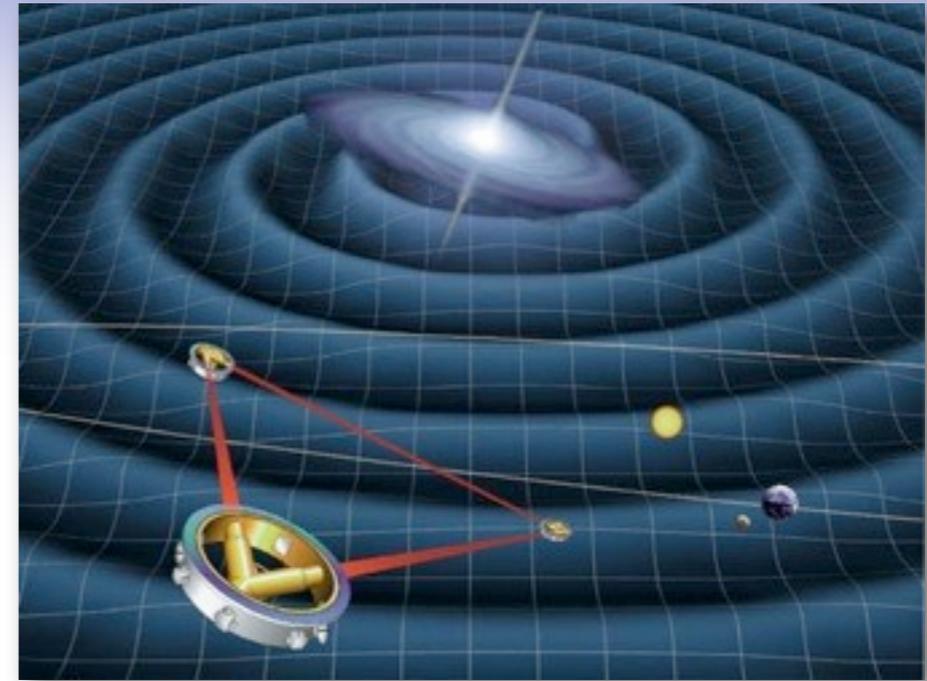
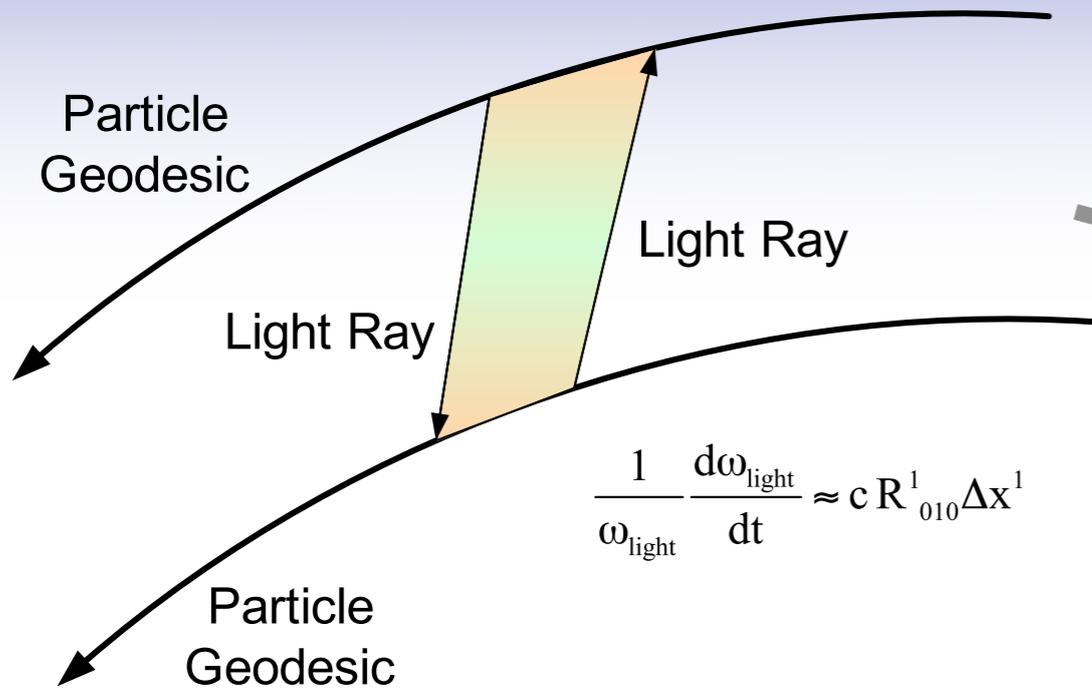
$$\frac{1}{\omega_{\text{light}}} \frac{d\omega_{\text{light}}}{dt} \approx c R^1_{010} \Delta x^1$$



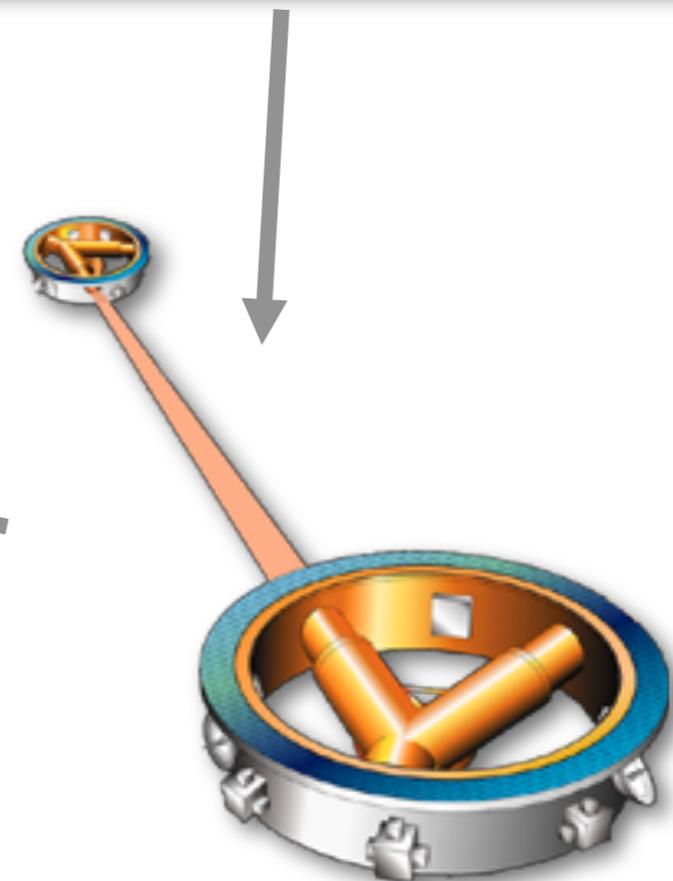
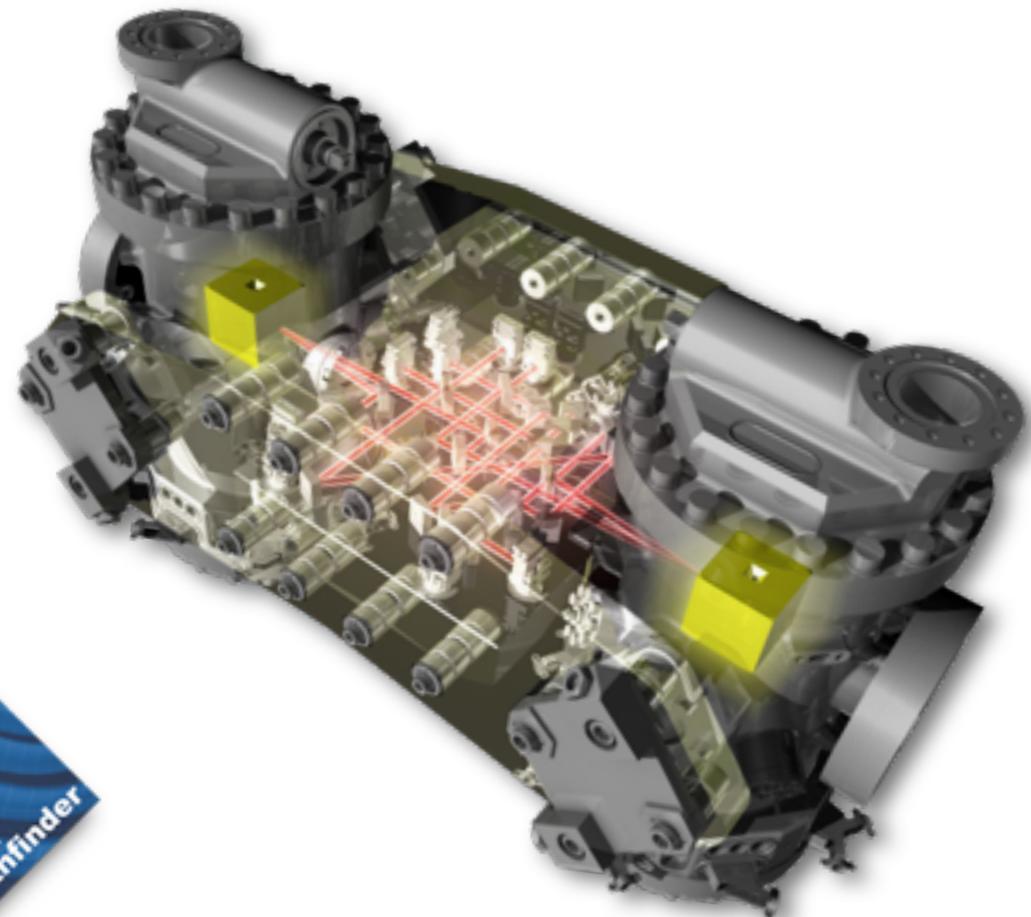
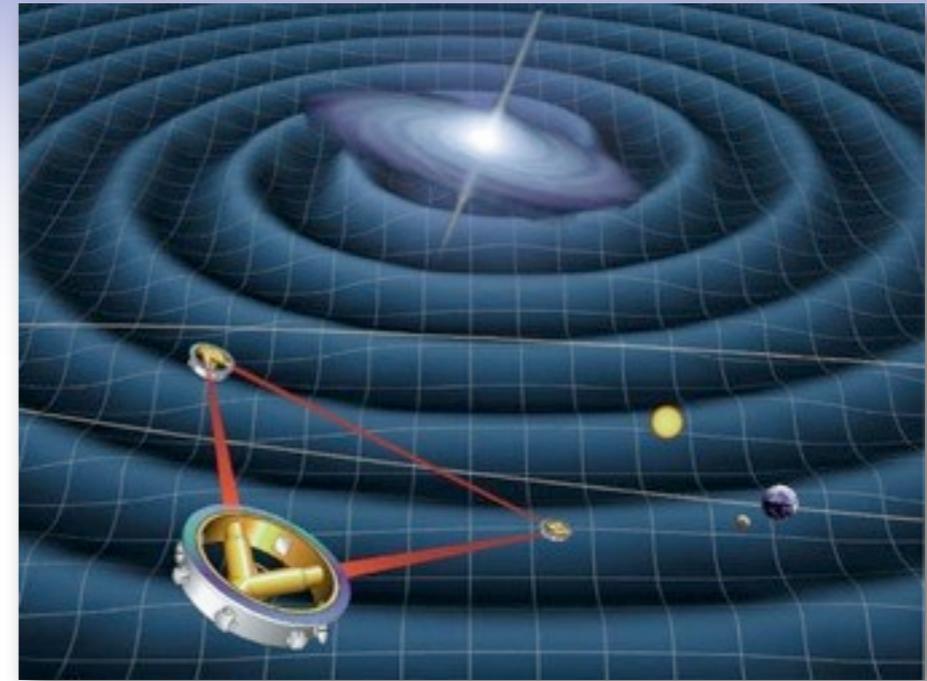
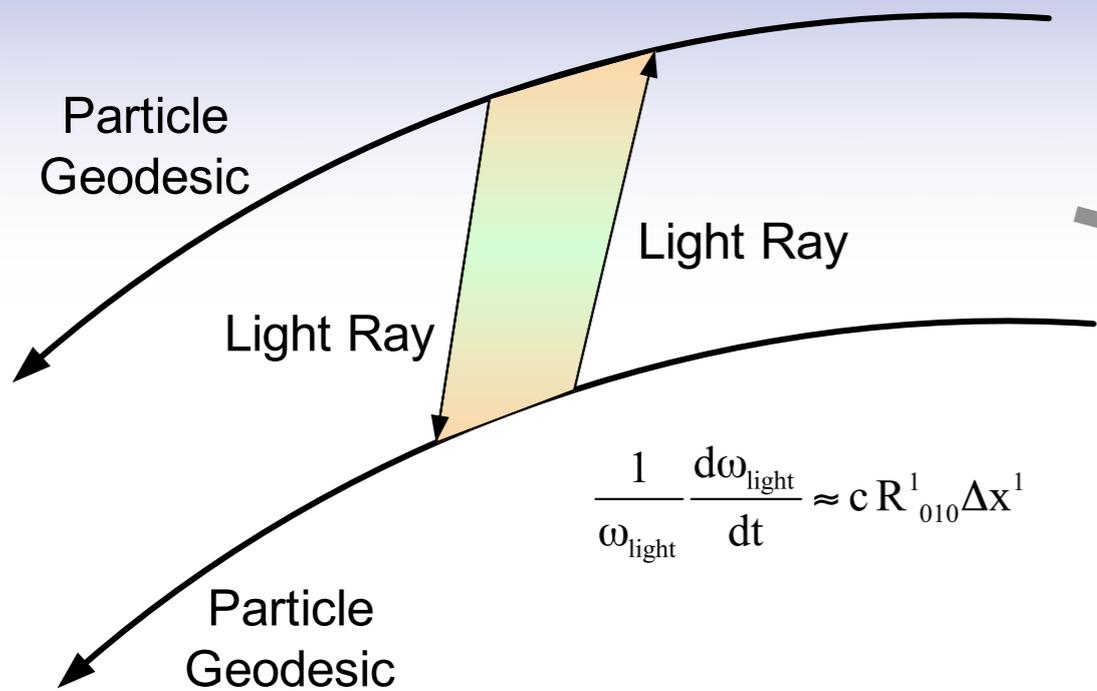
Mission Concept



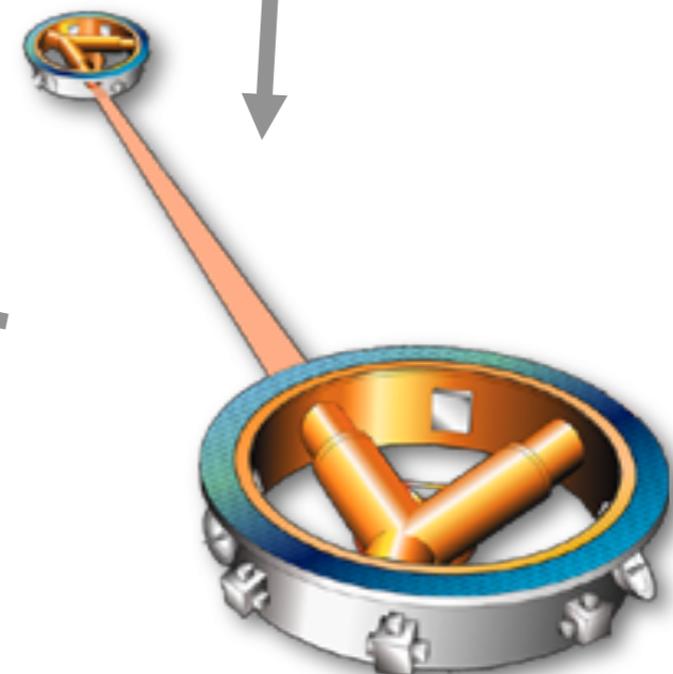
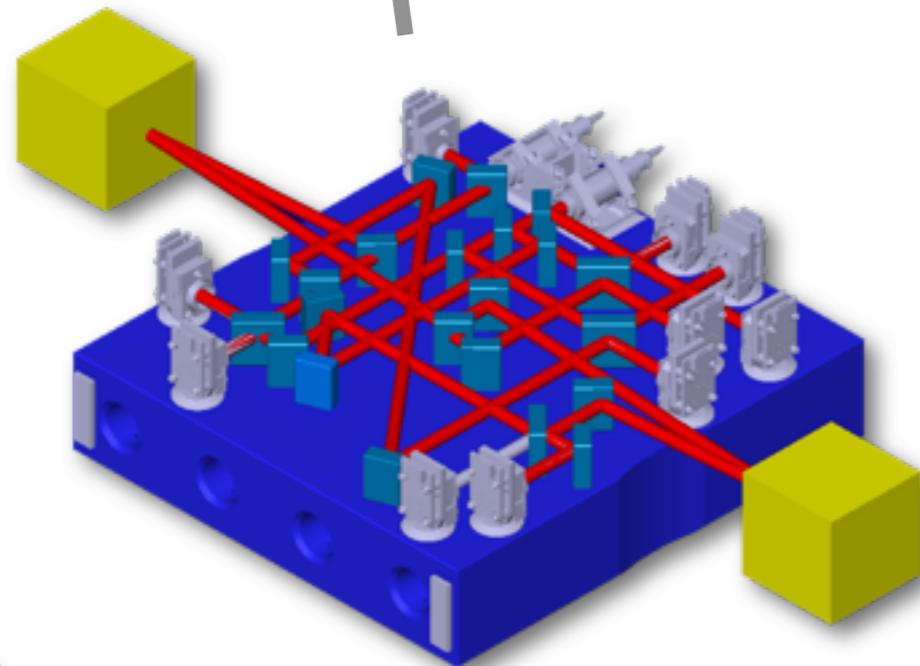
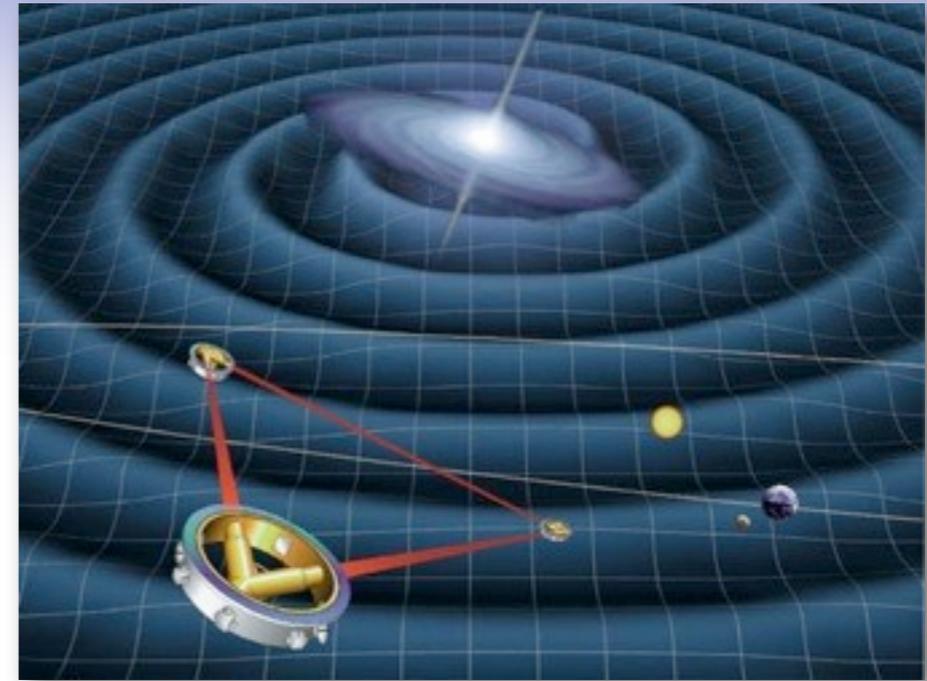
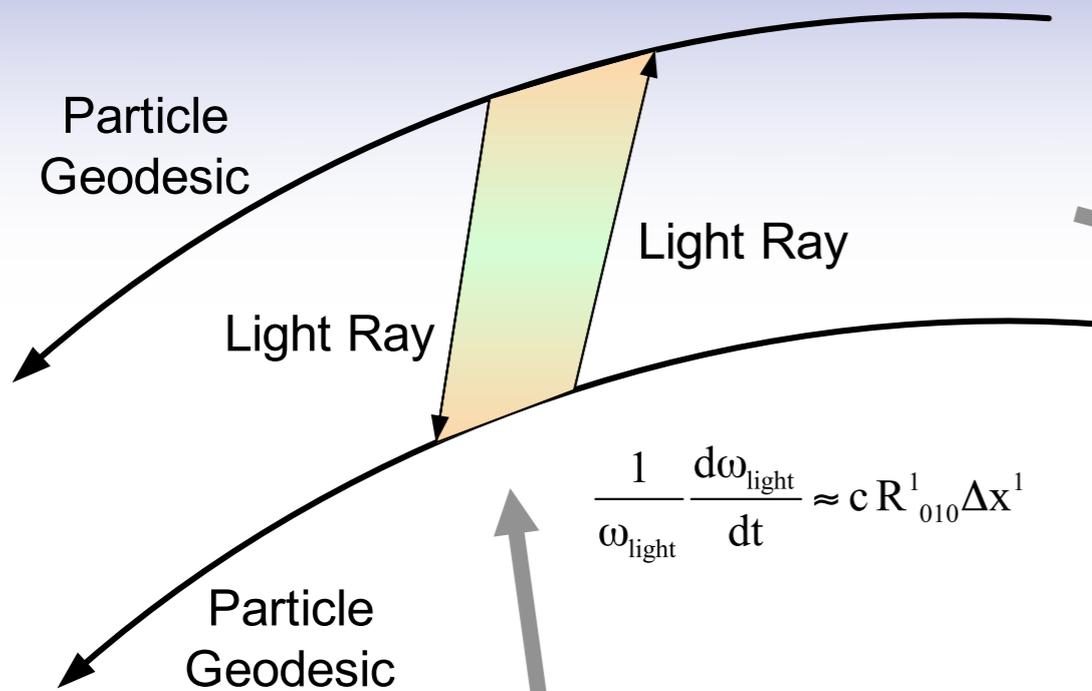
Mission Concept



Mission Concept



Mission Concept



LISA Pathfinder

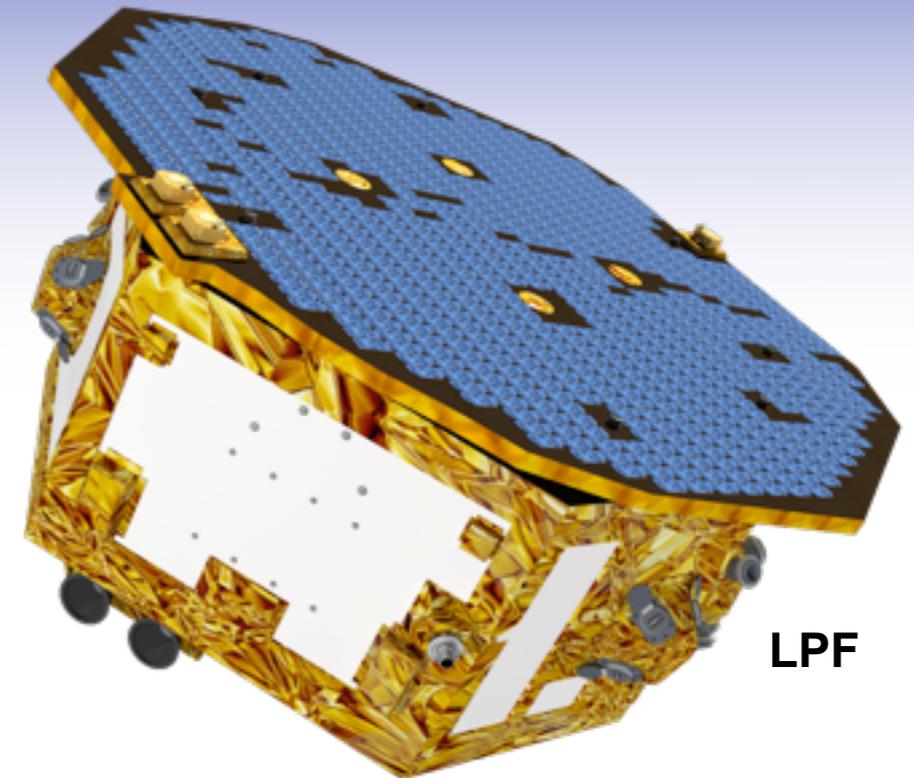
 LISA Pathfinder consists of:

– Spacecraft

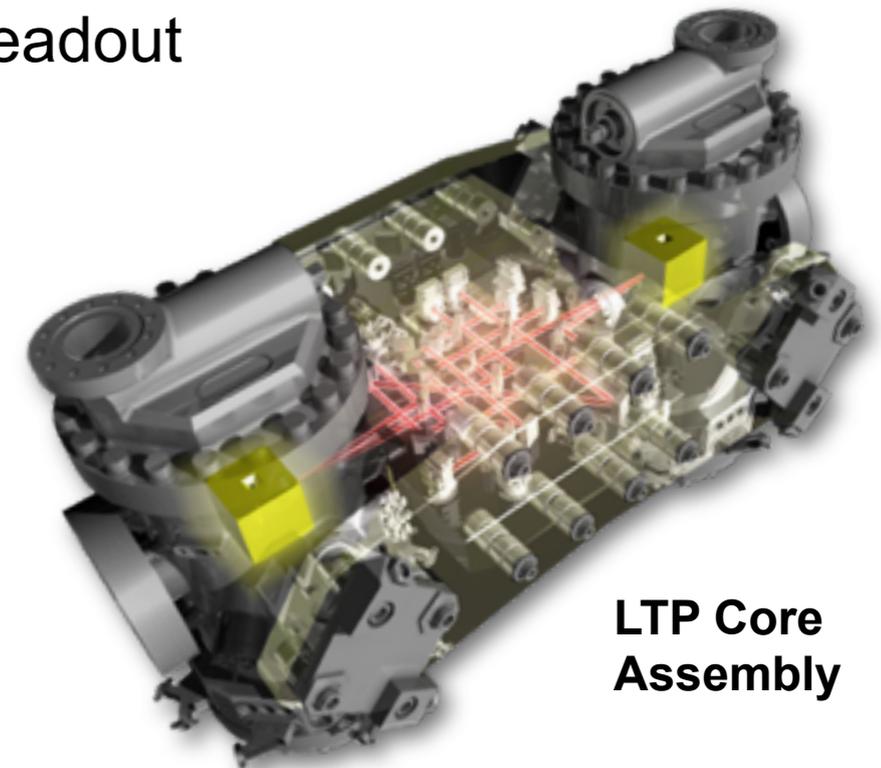
- Provided by ESA
 - Industrial Prime Contractor: Astrium UK
- s/c also includes the drag free control software (DFACS) and micro-Newton thrusters

– Payloads

- **The LISA Technology Package (LTP)**
 - Provided by European member states and ESA
 - Consists of inertial sensors, and interferometric readout
- **The Disturbance Reduction System (DRS)**
 - Provided by NASA
 - Consists of processor running drag-free control software and micro-Newton thrusters



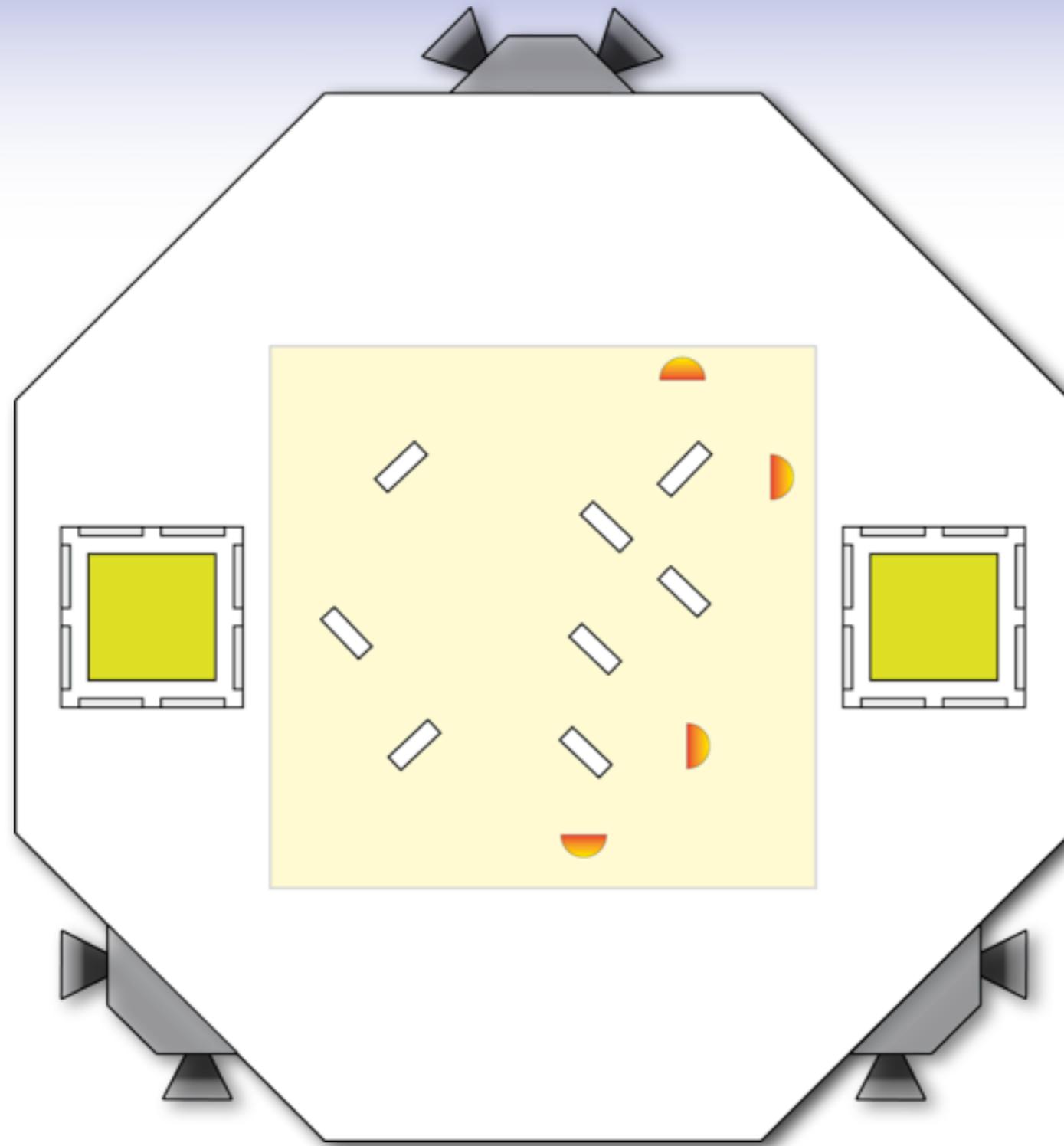
LPF



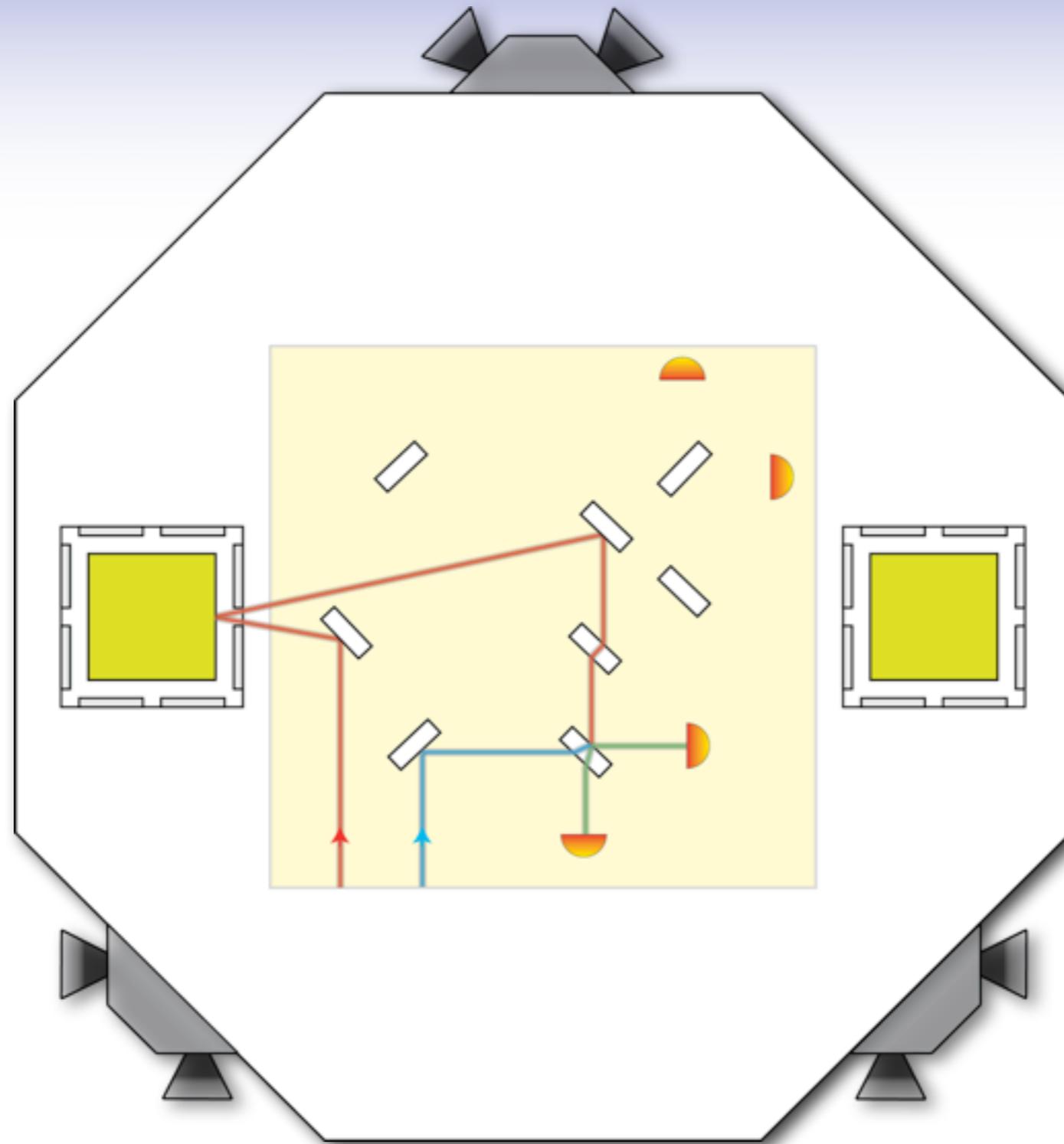
LTP Core Assembly



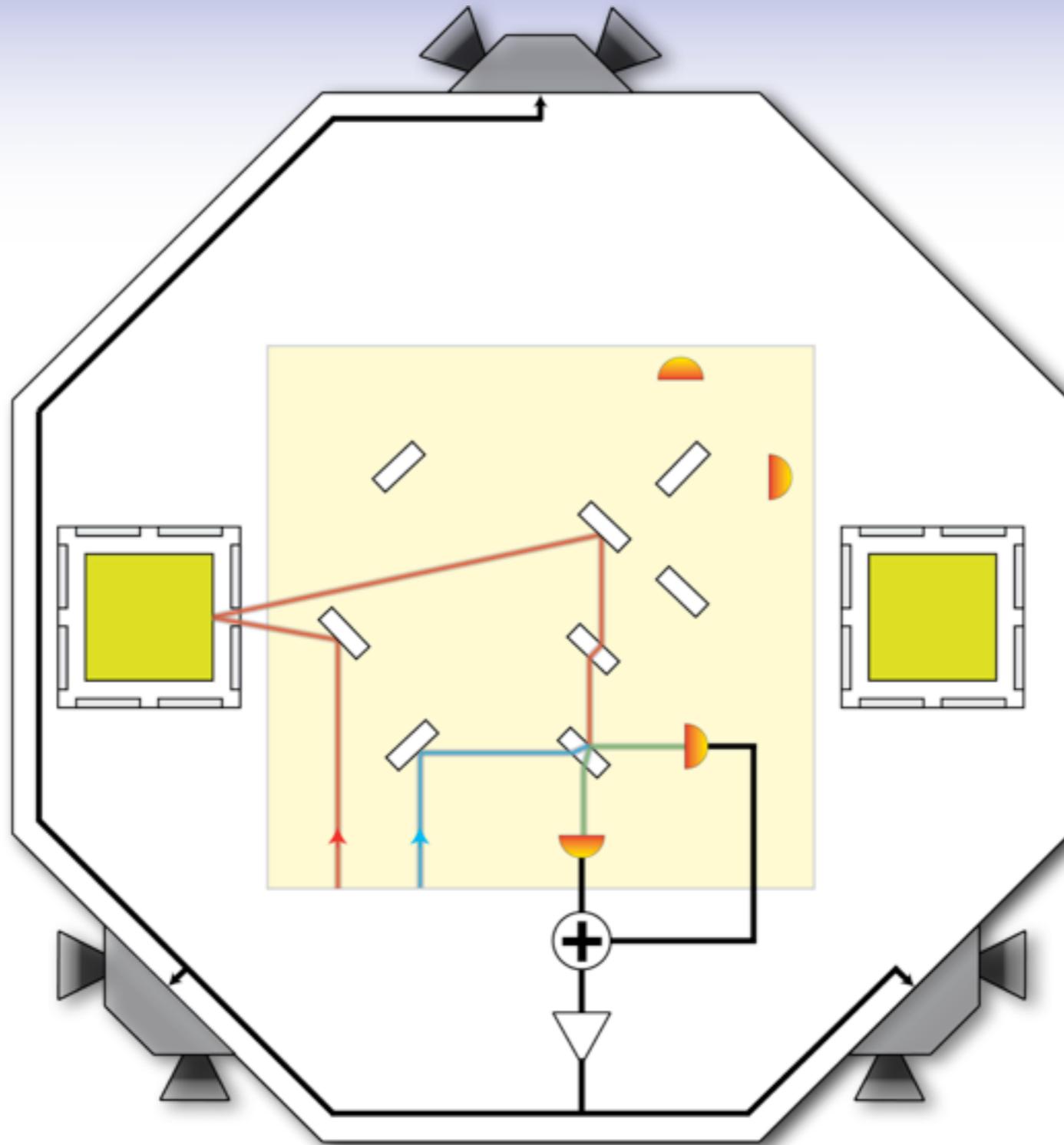
DFACS Principle



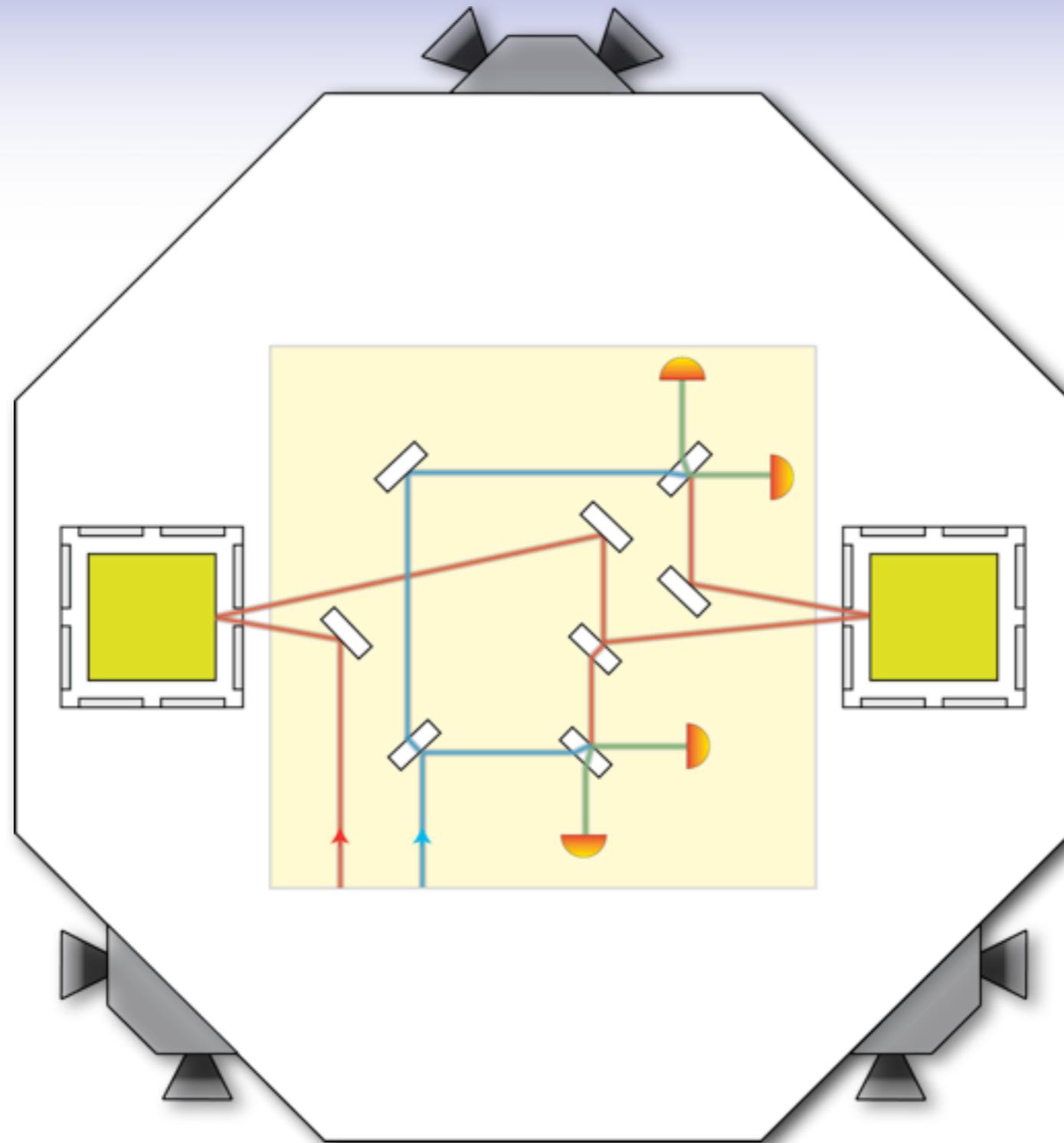
x_1 interferometer



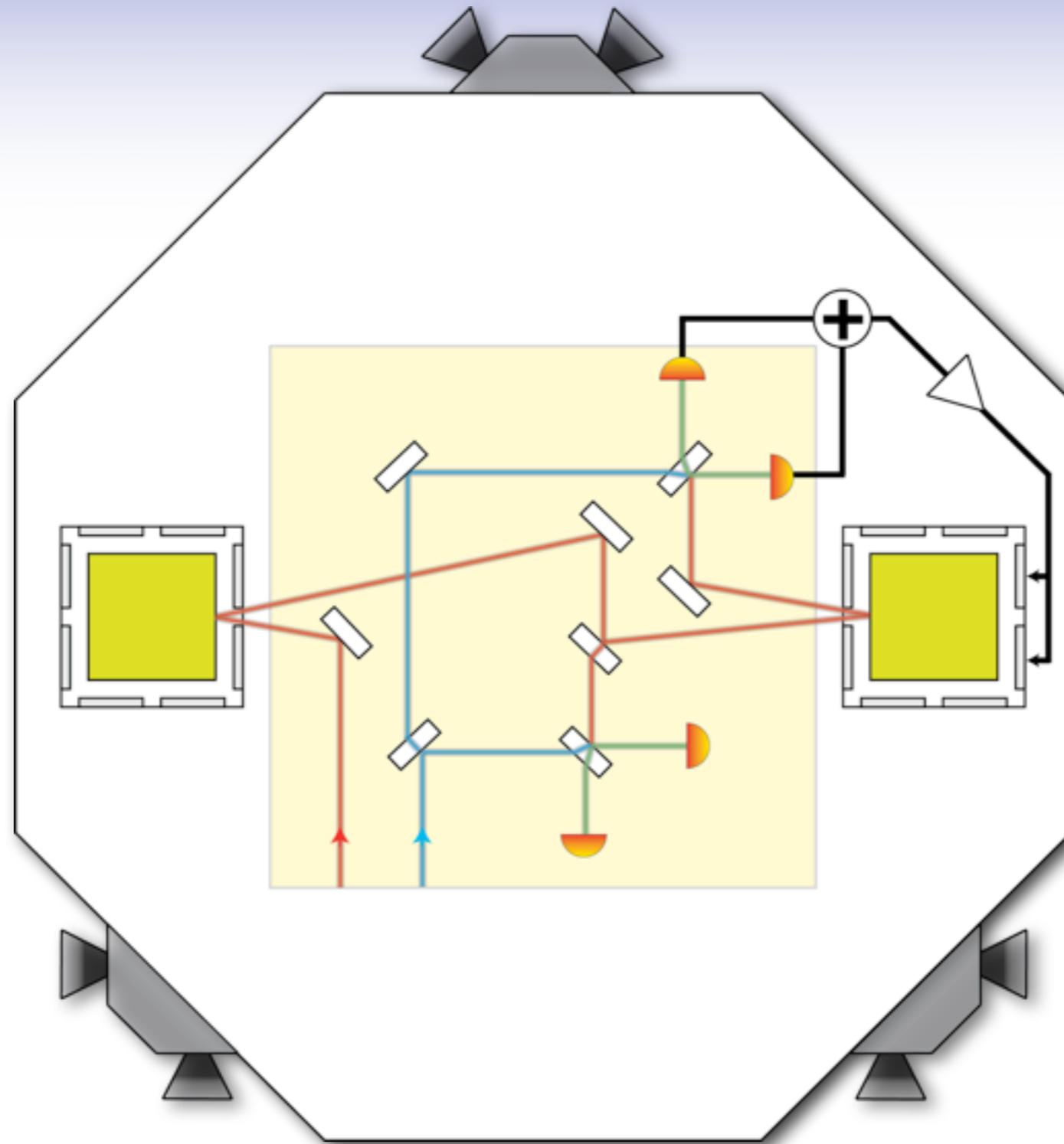
x_1 Drag Free



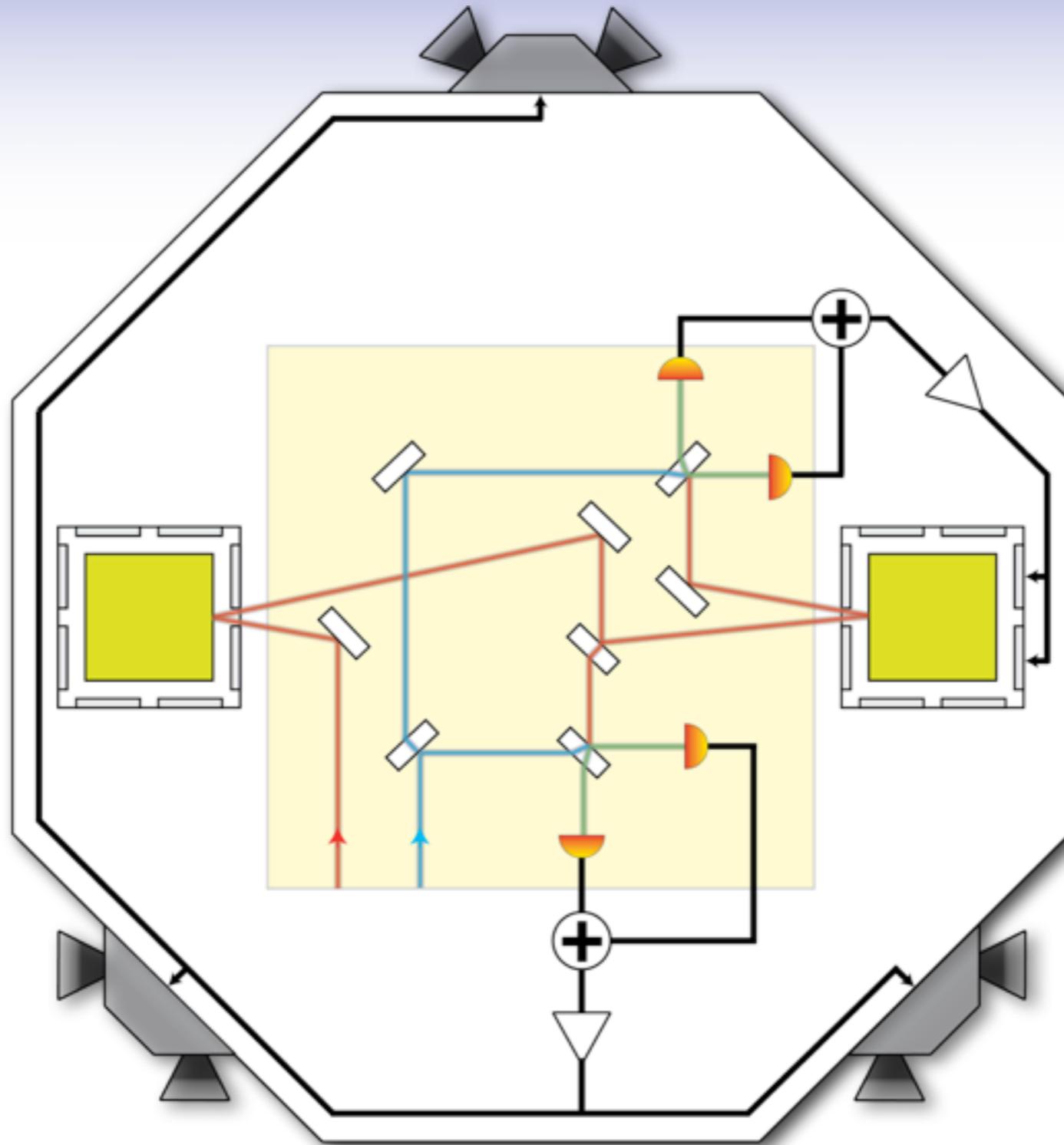
X_2-X_1 Interferometer



x2 low frequency control



DFACS



LPF Development Approach

-  LPF is a *Pathfinder* for spaceborne gravitational wave detectors
 - System design is for NGO
-  Relaxation of performance requirements is only allowed to make testing conditions achievable. It shall not affect the design.
-  Implications:
 - When one subsystem reaches a given level of readiness for LISA Pathfinder (CDR, QR, FAR, etc) it reaches the same level of readiness for NGO too.
 - All the development risks up to that stage are consequently retired also for NGO
 - Risk of unforeseen system level problems (estimated to be low) is finally retired by the flight.



NGO basic design items

Free flying test mass subject to very low parasitic forces:

- Drag free control of spacecraft (non-contacting spacecraft)
- Low noise microthruster to implement drag-free
- Large gaps, heavy masses with caging mechanism
- High stability electrical actuation on cross degrees of freedom
- Non contacting discharging of test-masses
- High thermo-mechanical stability of S/C
- Gravitational field cancellation

Precision interferometric, *local* ranging of test-mass and spacecraft:

- pm resolution ranging, sub-mrad alignments
- High stability monolithic optical assemblies

Precision 1 million km spacecraft to spacecraft precision ranging:

- High stability telescopes
- High accuracy phase-meter
- High accuracy frequency stabilization
- Constellation acquisition
- Precision attitude control of S/C

Items that can only be validated in flight

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Items benefiting from in-flight validation

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NGO Design Items Validated By LISA Pathfinder

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Precision interferometric, *local* ranging of test-mass and spacecraft:

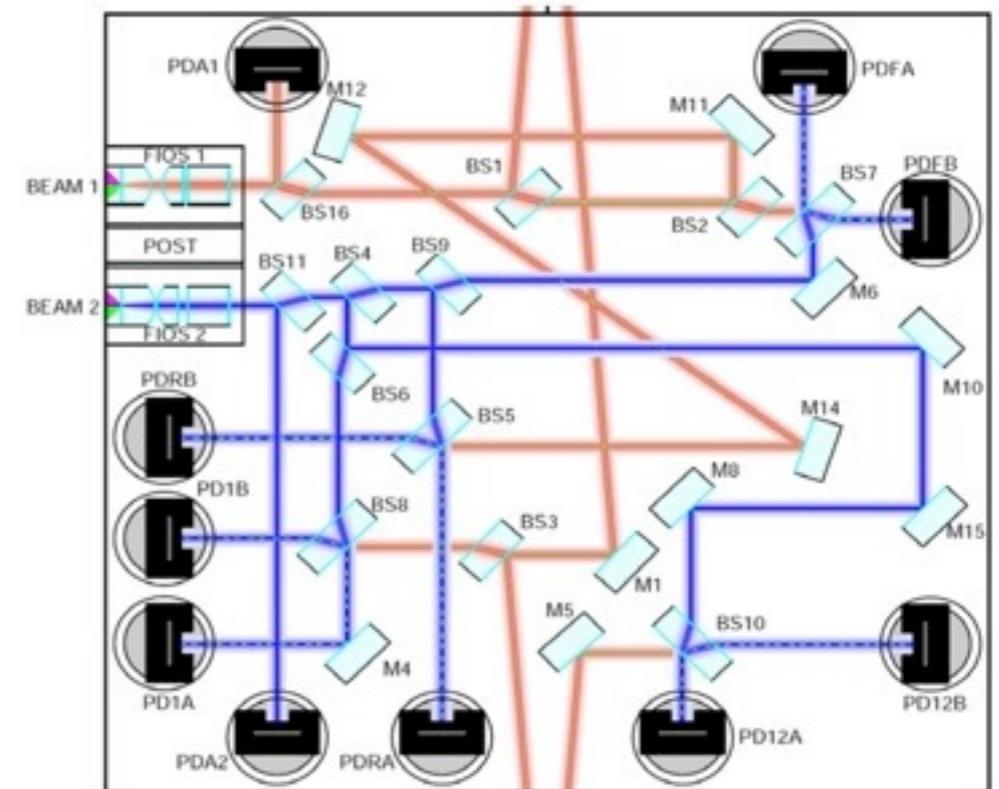
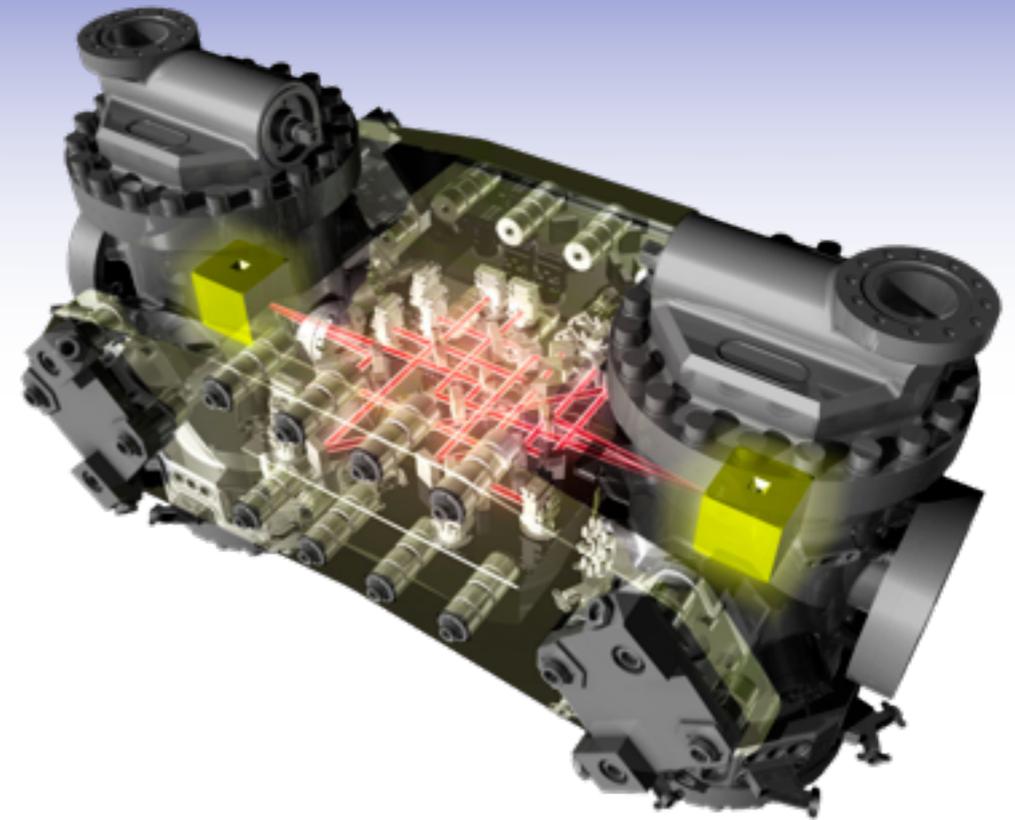
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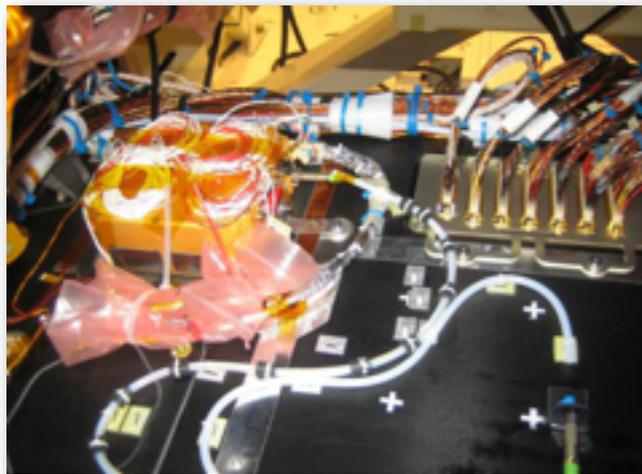
LTP Salient Features

- 🚀 LISA Technology Package (LTP) is the European instrument payload of LPF
- 🚀 Two Au:Pt test masses housed in separate vacuum enclosures
- 🚀 Relative position of test masses read-out by:
 - Heterodyne laser interferometry on sensitive axis
 - Capacitive sensing on all axes
- 🚀 Four interferometers on ultra-low expansion optical bench
 - x1, x2-x1, Frequency noise, reference interferometer

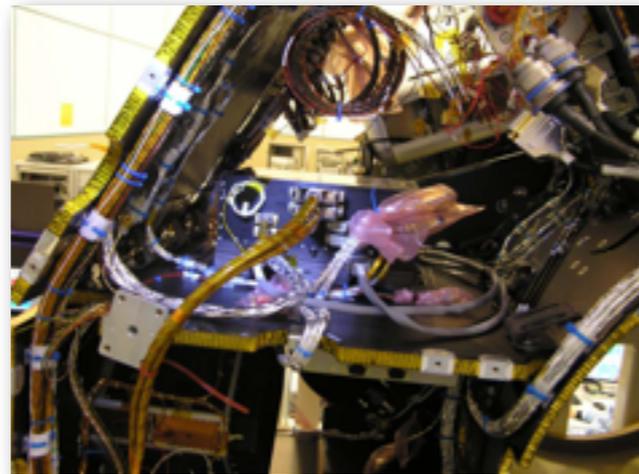


LTP Status

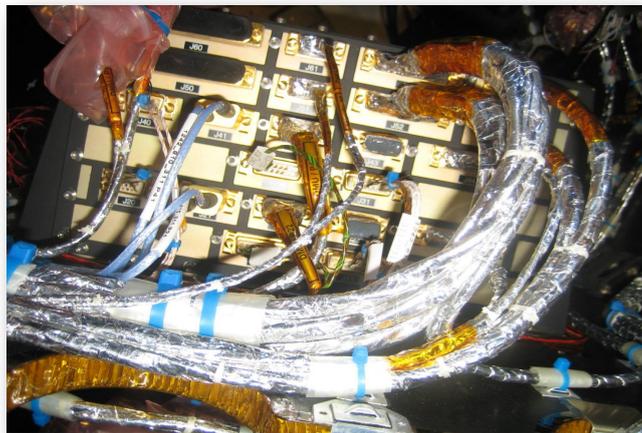
- 🚀 All electronic units have been delivered and integrated to the spacecraft
 - Laser, laser modulator, phasemeter, payload computer, UV lamp unit, ISS front-end electronics, diagnostics



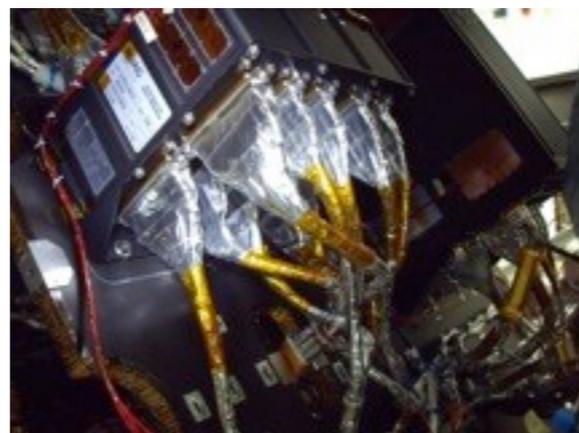
Reference Laser Unit



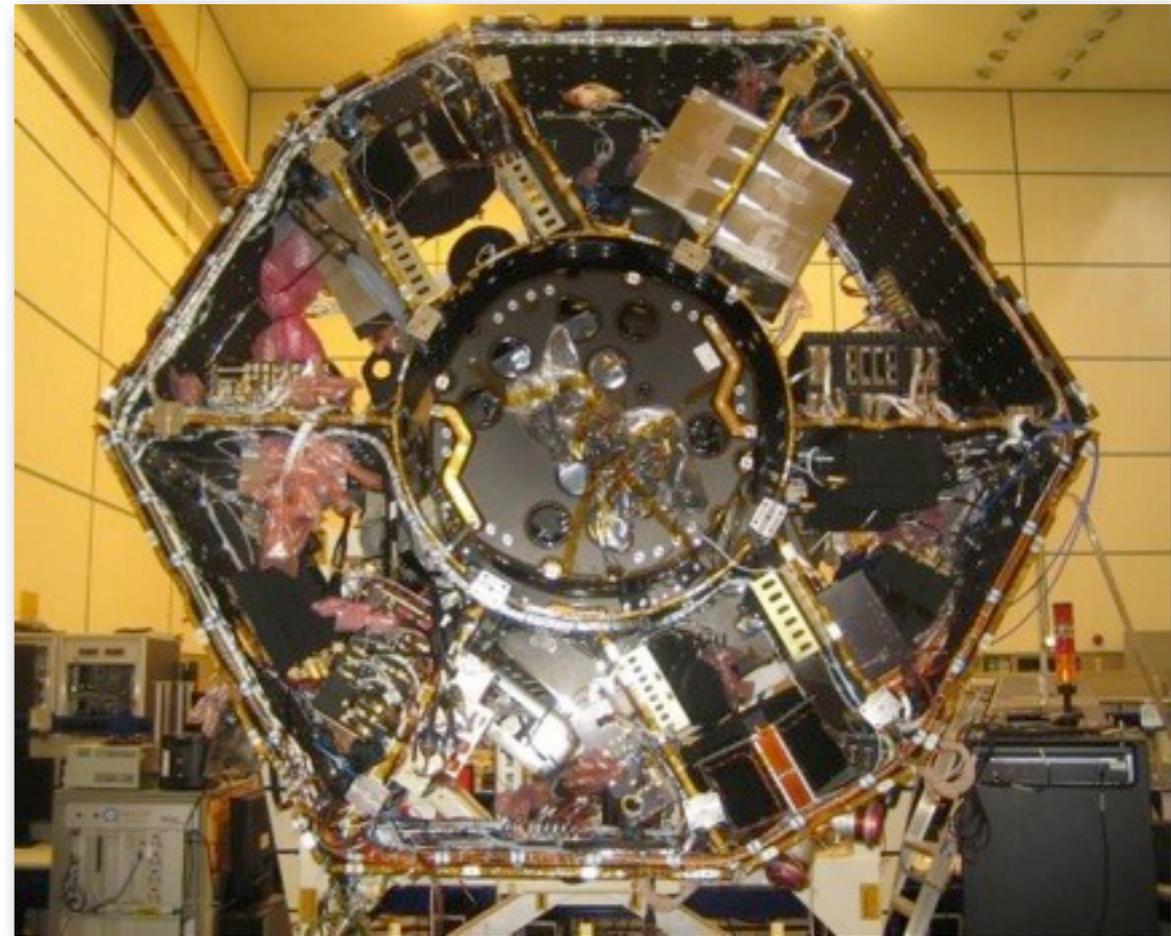
Phasemeter



Payload Computer



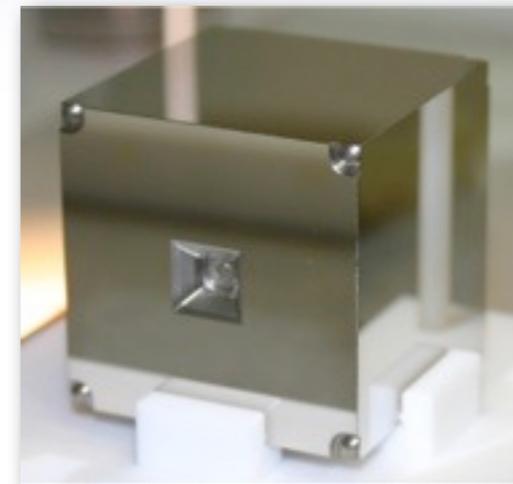
ISS Front-End Electronics



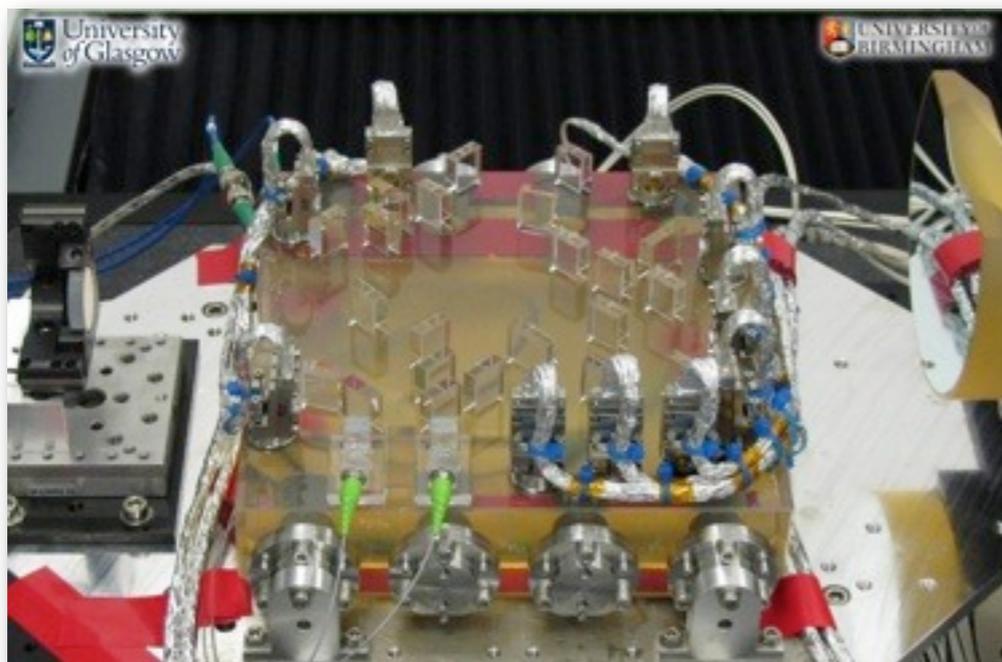
LPF during AIT at Astrium UK

LTP Status

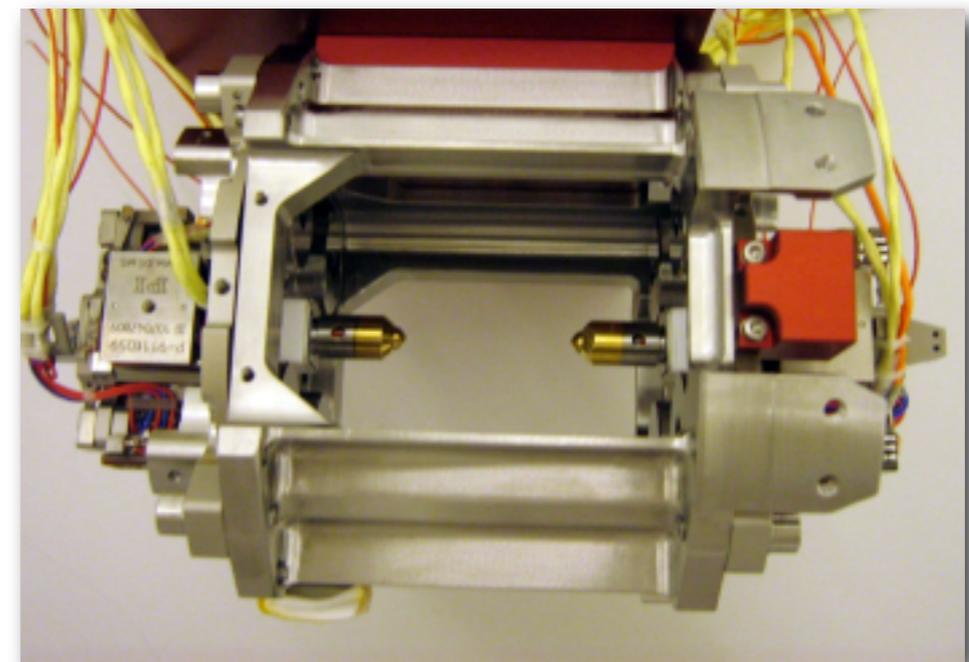
- Most of the components of the LTP Core Assembly are ready, however subsequent integration has been delayed.
 - Optical bench
 - Inertial sensor
 - Test Masses
 - Vacuum enclosure
 - Grabbing, positioning and release mechanism



Uncoated Au:Pt Test Mass



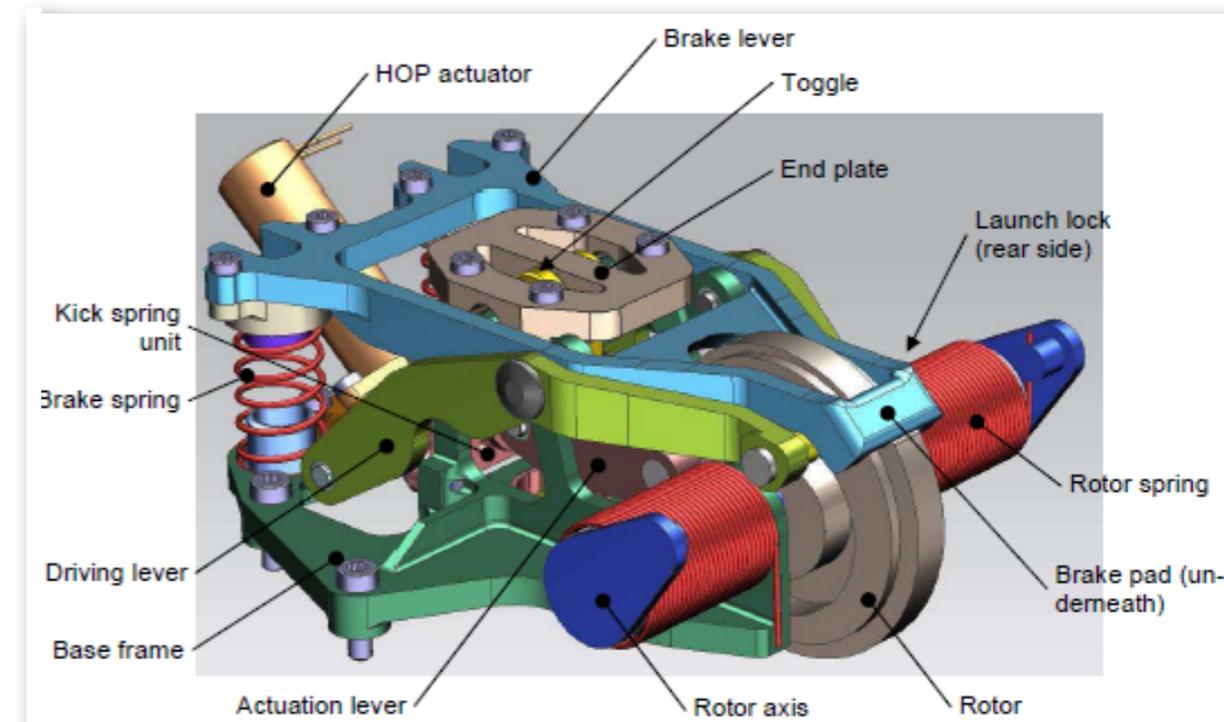
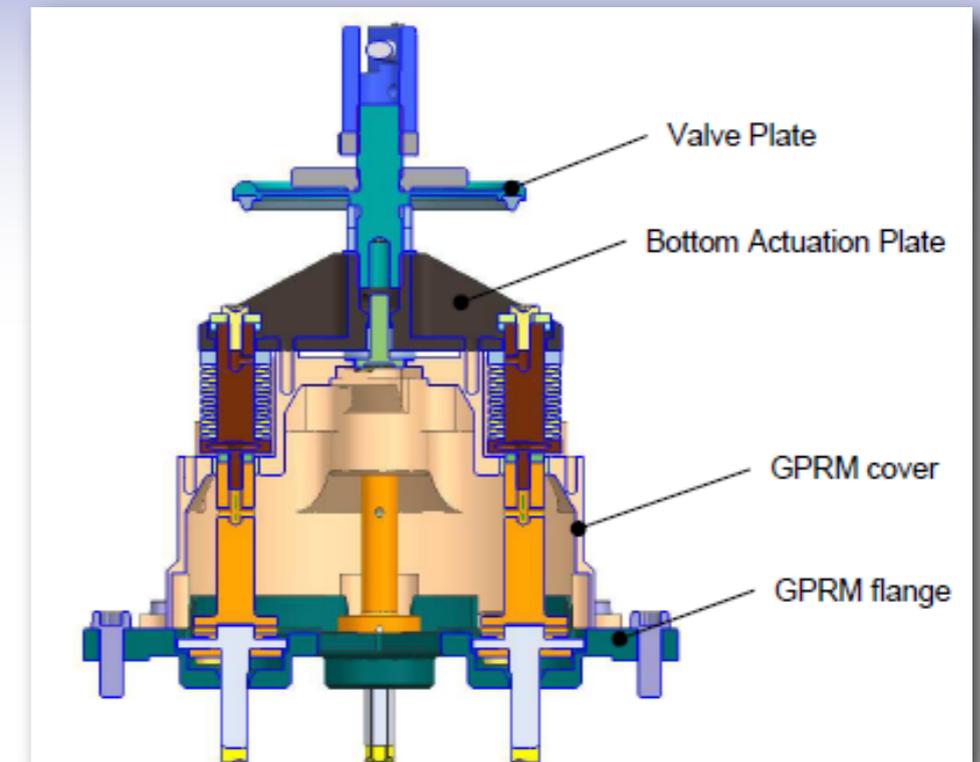
Optical Bench



Test Mass Grabbing, positioning and Release Mechanism (GPRM)

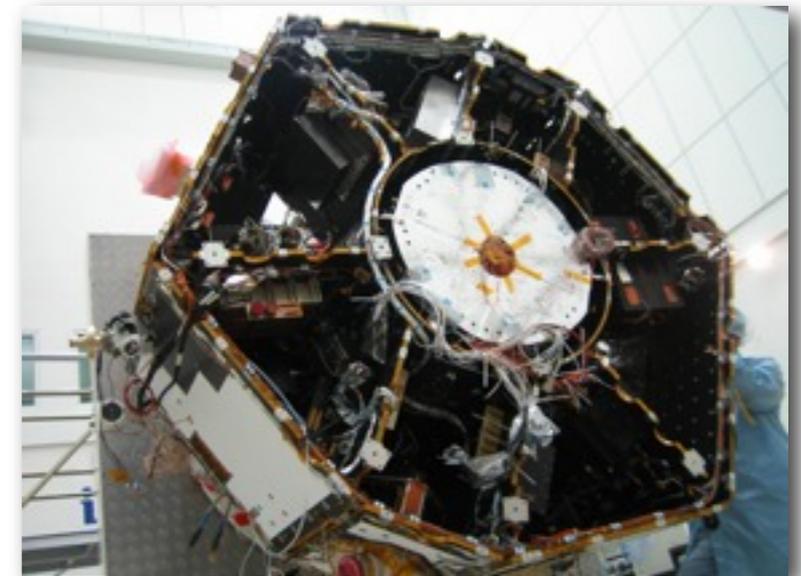
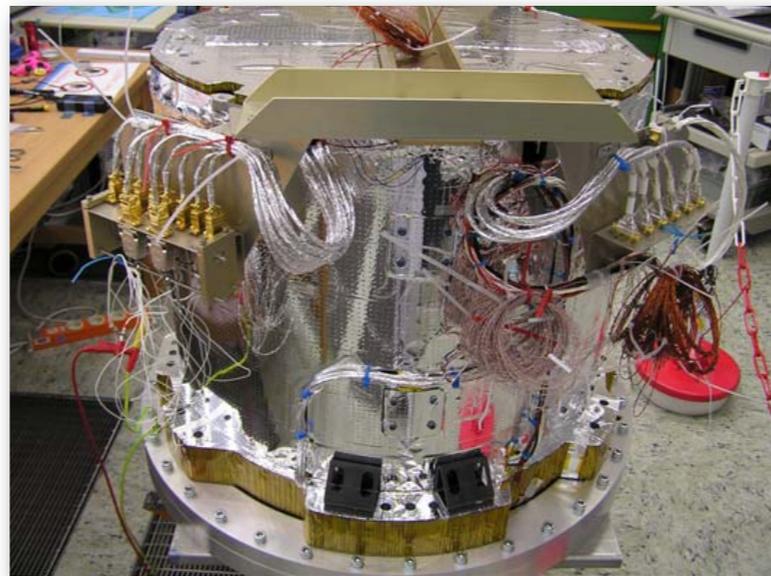
LTP Status

- Delays in the integration, and delivery, of the LTP Core Assembly are due to problems with the test mass launch lock
 - Original (hydraulic) launch lock failed during testing
 - A new design has now been selected, and flight units are being manufactured
 - New design is a much simpler, single-shot mechanism utilising a paraffin actuator
 - Delivery of FMs is scheduled for September 2012



TOQM

- ❖ In-lieu of the LCA, a *Thermal-Optical Qualification Model* (TOQM) was created to allow the environmental testing of the spacecraft to proceed
 - Consists of a flight optical bench, with flight mounting struts, and thermal mass dummies of the inertial sensors



LPF Spacecraft Bus

- ✈ All units of the spacecraft bus have been delivered and integrated
 - Only units not yet integrated are the micro-thrusters (which are mounted on the outside on the spacecraft)
- ✈ System environmental tests are now complete
- ✈ Following closed-loop tests in January, spacecraft will be put into storage awaiting the delivery of the LCA

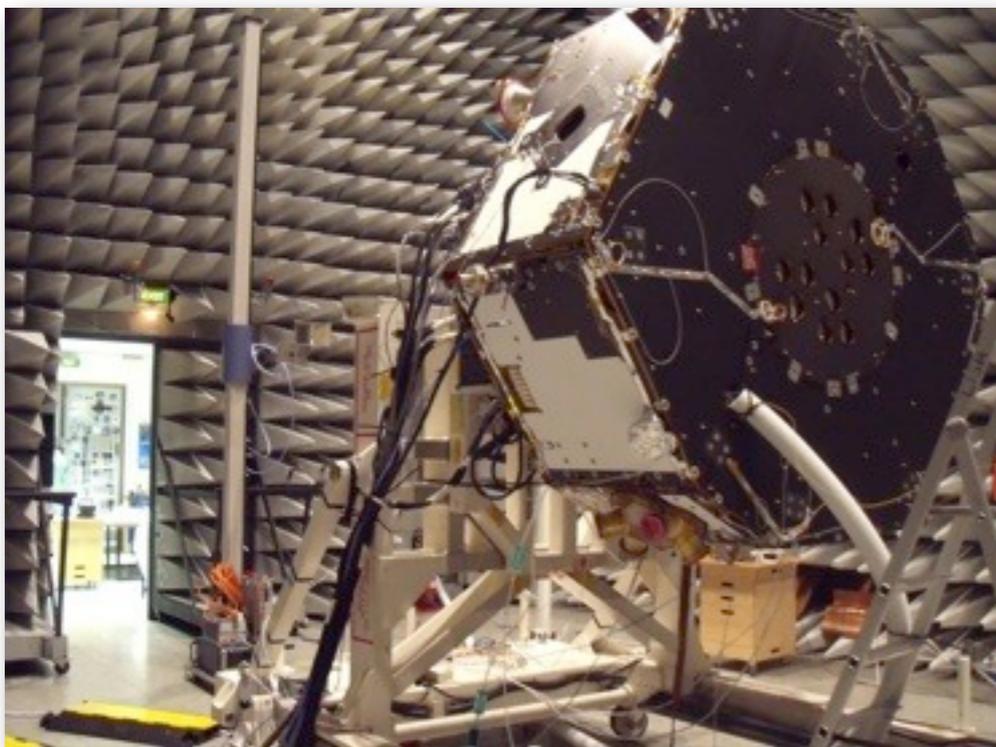
System Test Campaigns



Sine Vibration



Transfer Orbit Thermal Test

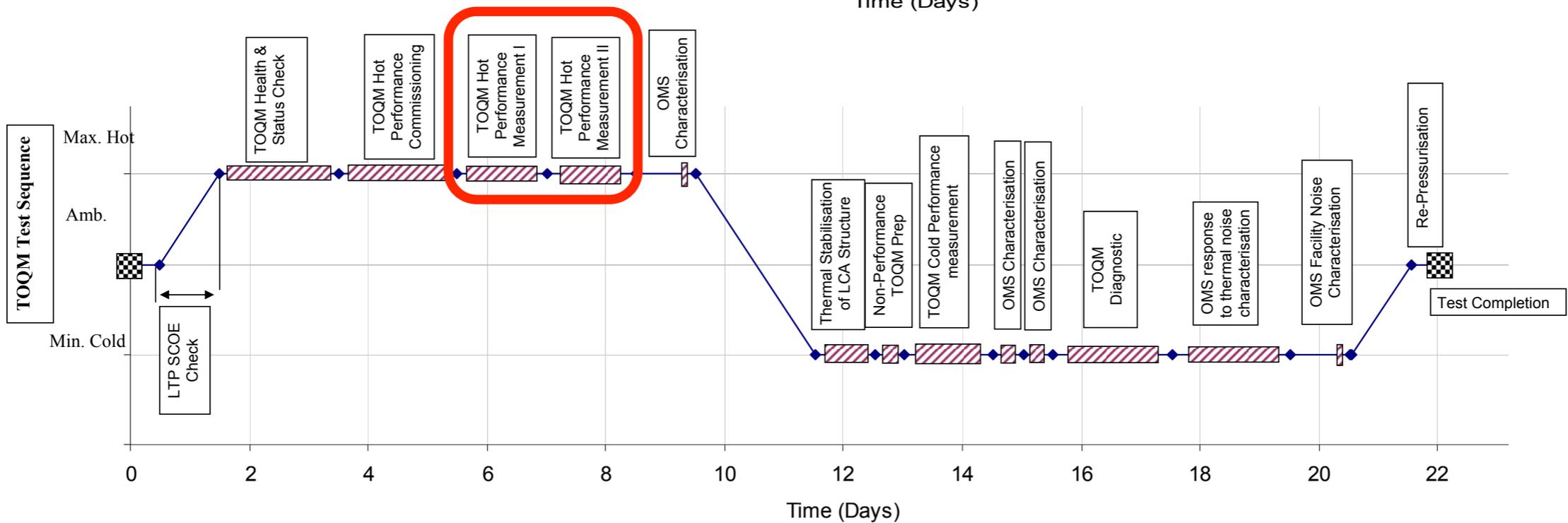
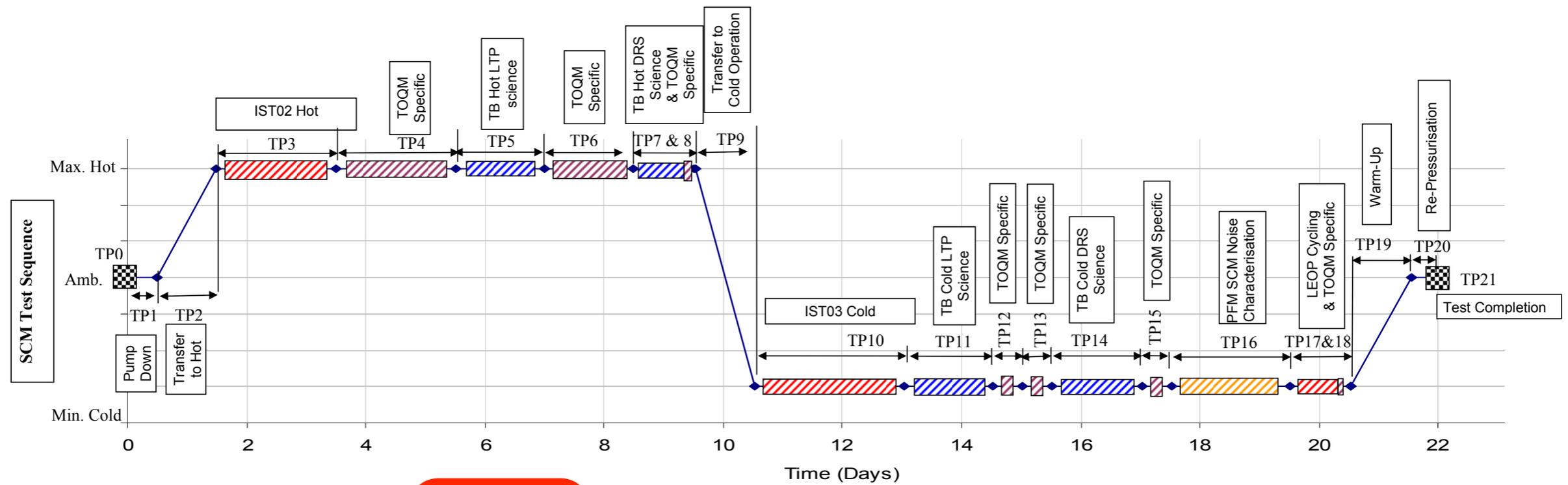


EMC

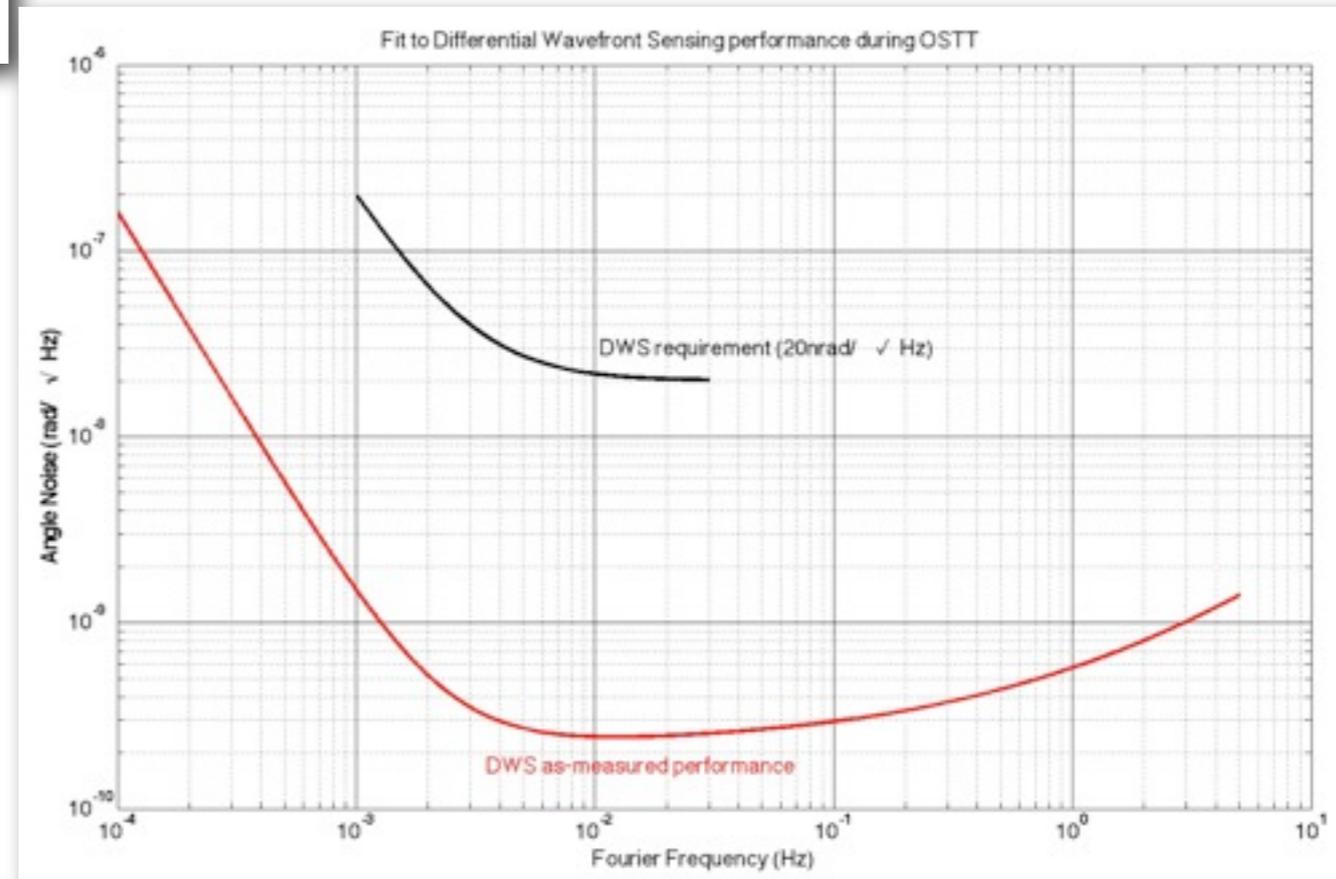
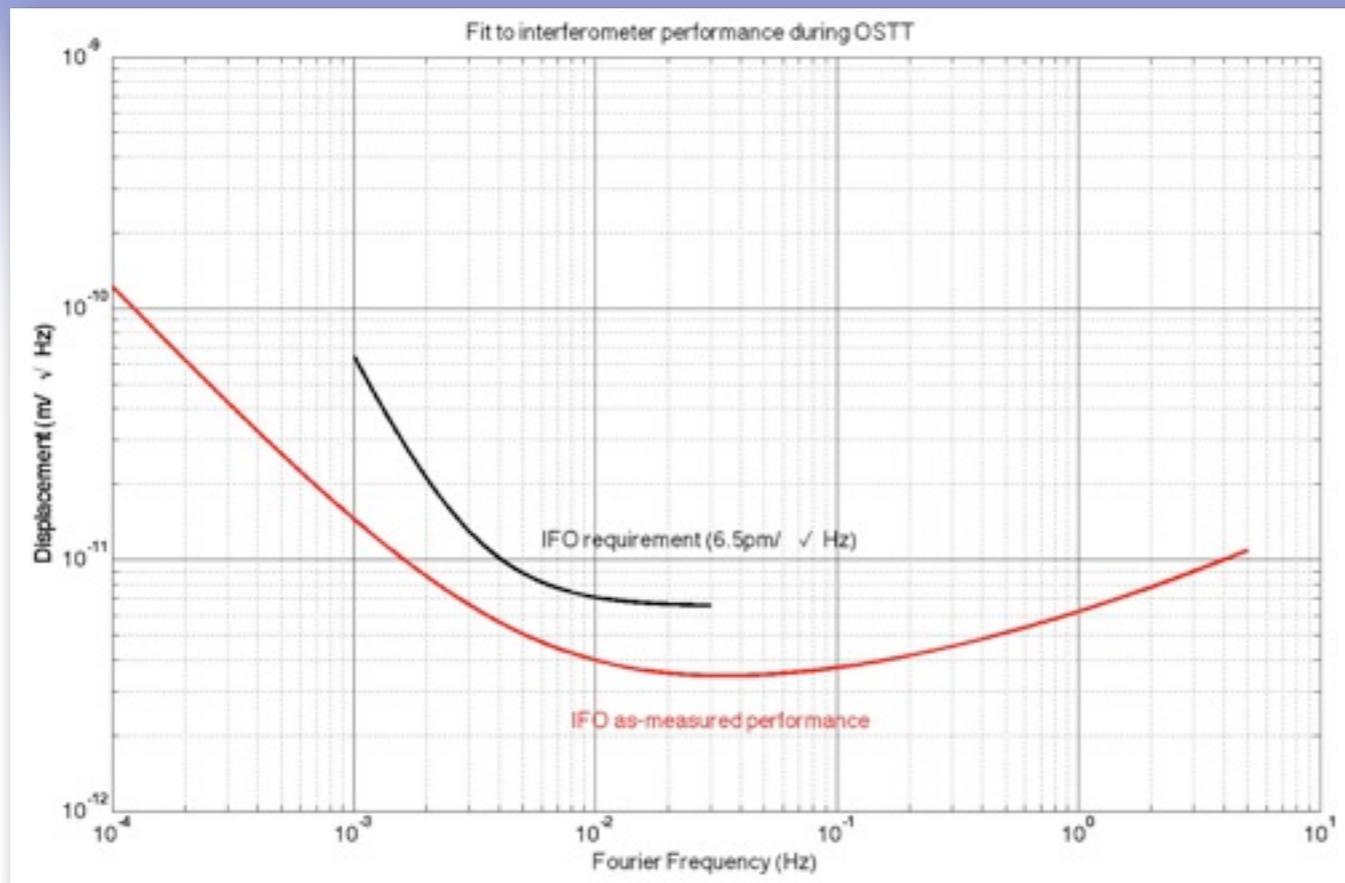


On-Station Thermal Test

OSTT Sequence



Fit to OSTT Performance



Micro-thrusters

-  LPF will carry two sets of micro-thrusters
 - The NASA supplied Colloidal Thrusters
 - European supplied thrusters
-  The baseline thruster for LPF is the Cs Field Emission Electric Propulsion (FEEP) thruster
-  Due to delays in the qualification schedule of the FEEPs, the project has recently instigated a review of the use alternative thrusters for LPF
 - Cold Gas thrusters (the GAIA thruster)
 - Micro Radio-frequency Ionisation Thrusters (μ -RIT)

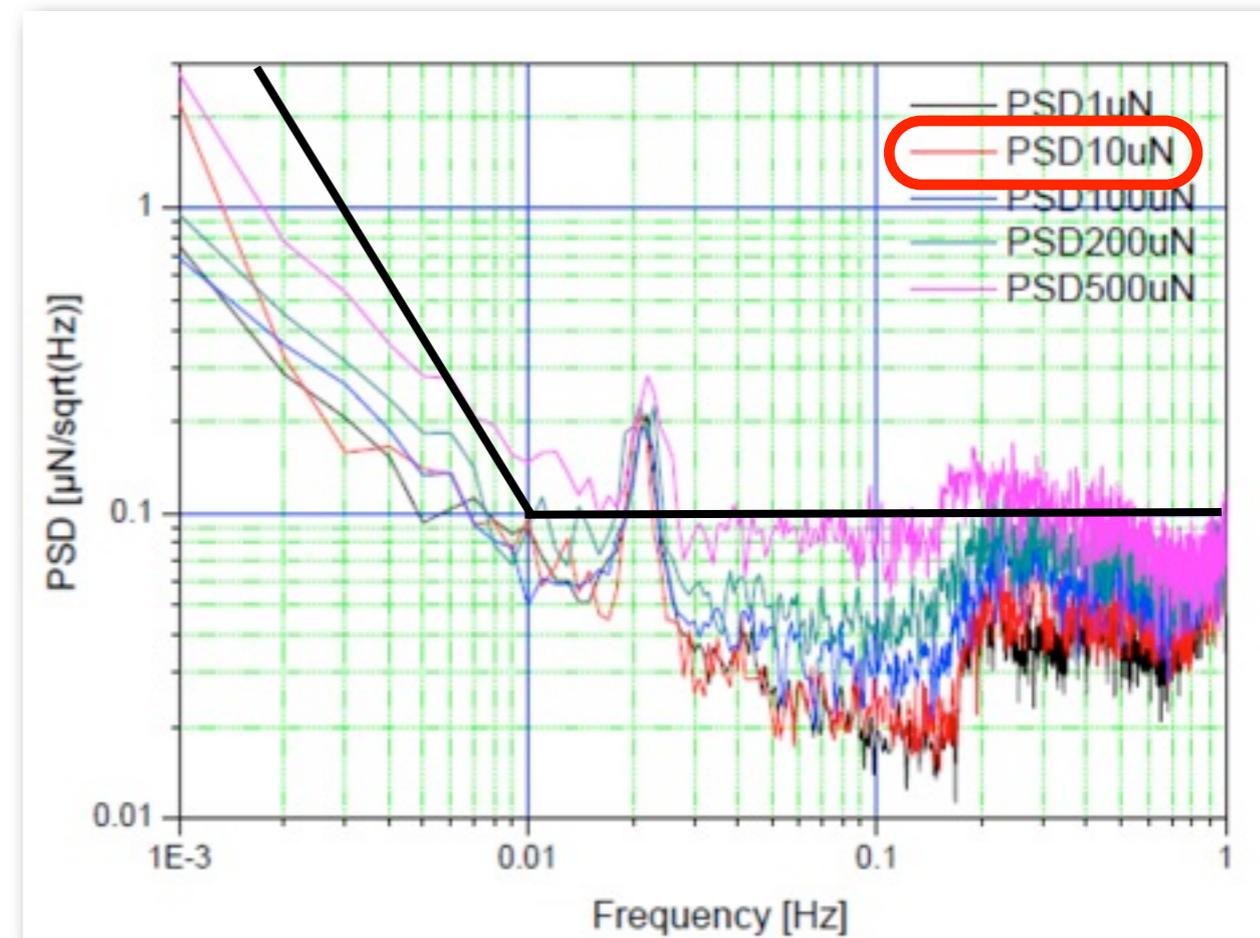
Micro-Thrusters

Review Board Conclusions (May 2011):

- Continue FEEP development, culminating in two tests demonstrating the LPF lifetime (600Ns) by June 2012
- Industry to study the cold gas thrusters, with PDR to be held in October 2011
- Industry to study micro-RIT thrusters with SRR in September 2011, and PDR by the end of the year

As of now:

- Cold Gas PDR was successful
 - Major issue is the schedule for the delivery of the thruster units
 - Board recommended to start the procurement of the long-lead items before the decision is made on the flight thruster technology
- Micro-RIT
 - SRR has been held
 - Development of the technology will continue up to PDR level



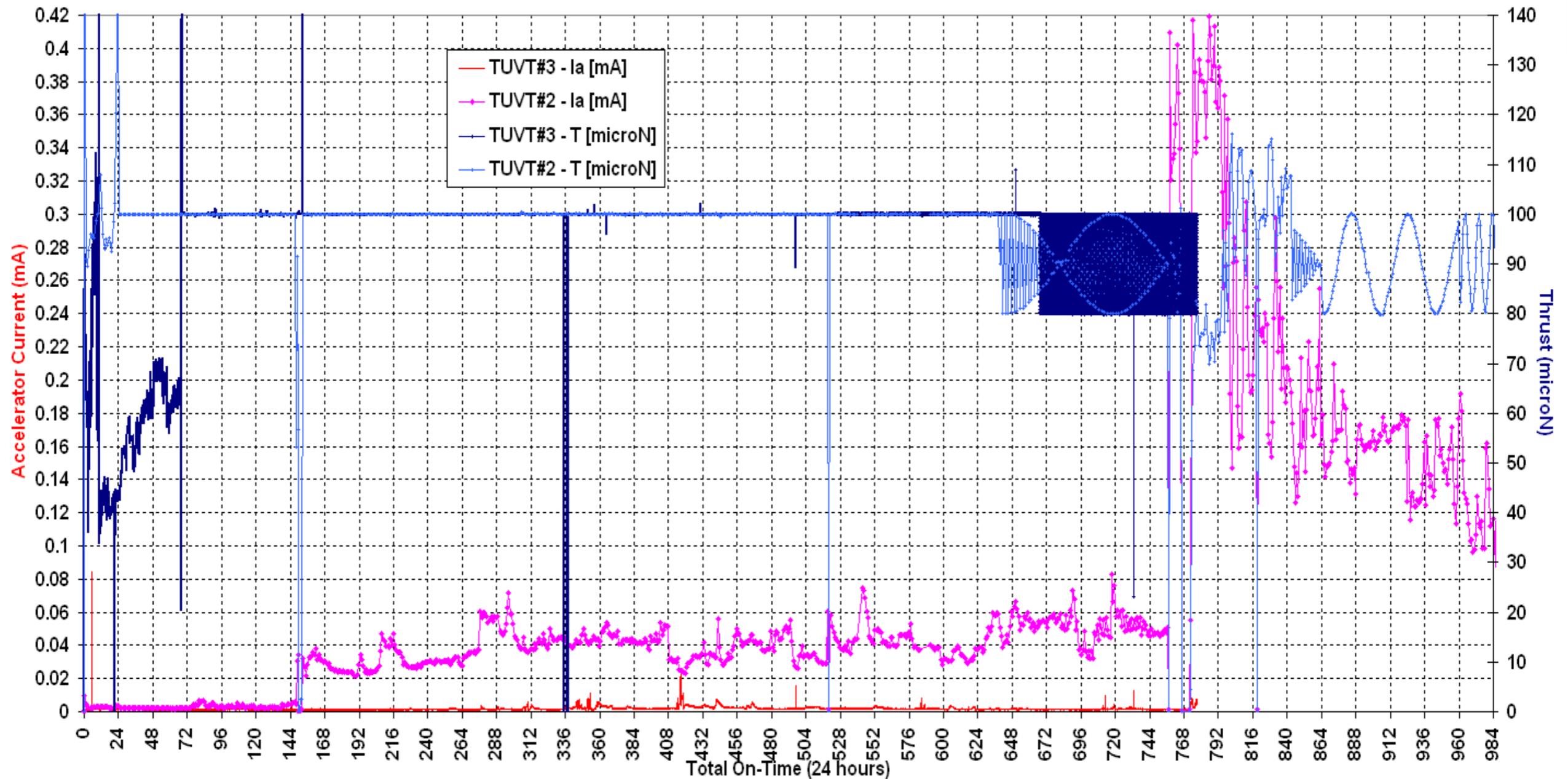
Measured thrust noise of GAIA thrusters

Micro-Thrusters

- 🚀 Review Board requested that the FEEP tests continue into next year
 - Culminating in two thruster assembly endurance tests, both of which must demonstrate the full LPF total impulse requirements.
- 🚀 After thorough investigation, the FEEP thruster head has been redesigned to limit caesium leakage from the emitter slit
 - Achieved in part by narrowing the slit width from $1.3\mu\text{m}$ to $0.7\mu\text{m}$
- 🚀 Ongoing thruster unit validation test is demonstrating that the new design of thruster is working perfectly!!
- 🚀 Downselect on which thruster to take to flight will be made in June 2012.

Micro-Thrusters

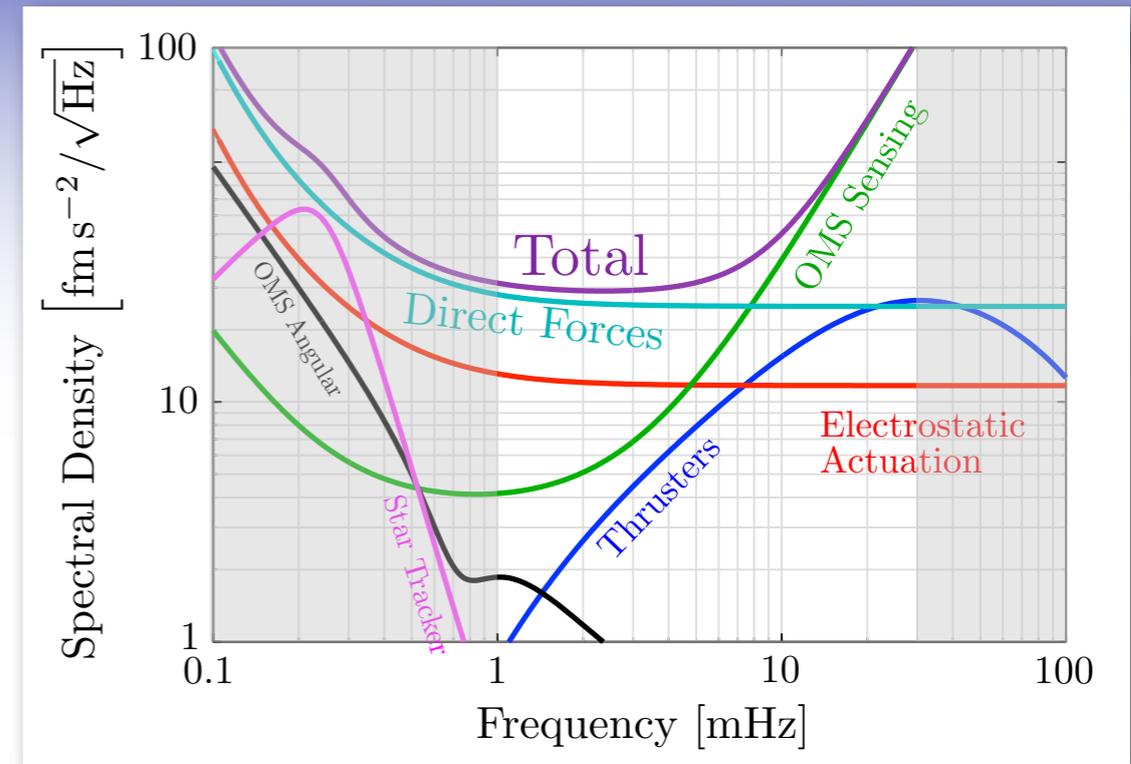
TUVT#3/TUVT#2 - Accelerator Current



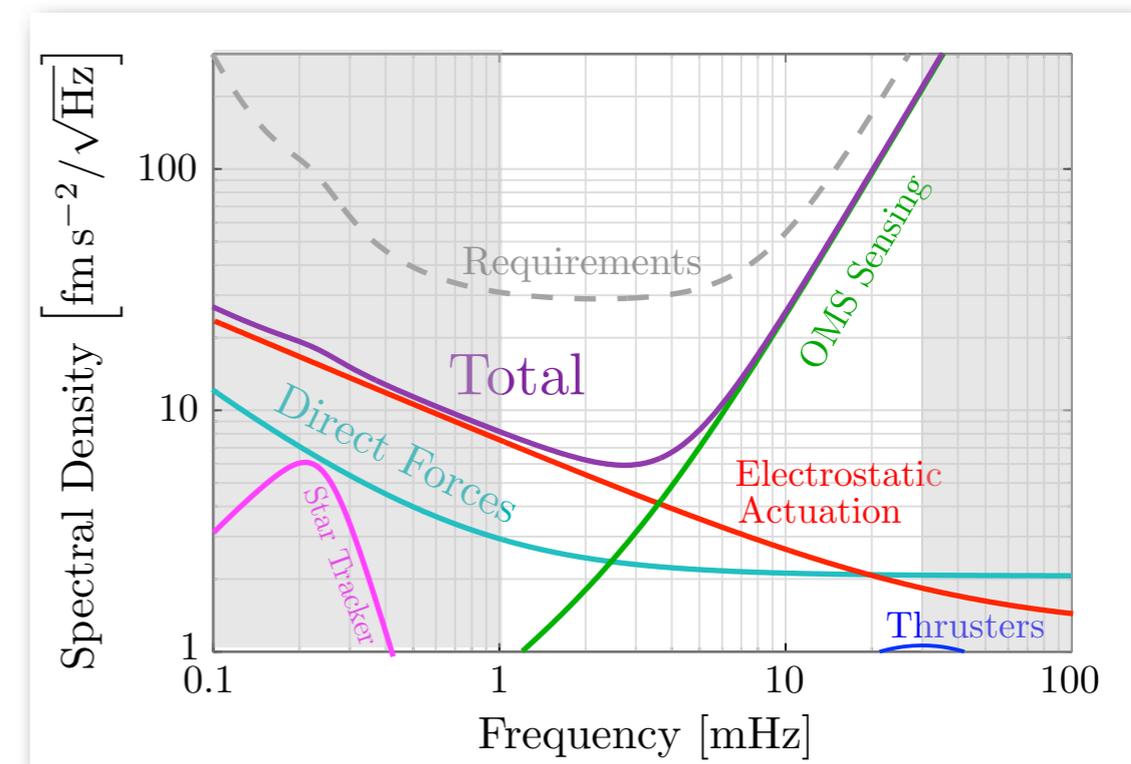
NB: As of today, total impulse now at ~300Ns (~900hours). Equivalent to ~50% of LPF total impulse requirement

LPF Performance

- Extensive performance noise model has been developed by both the PI and industry
- The main goal of LPF is to validate this noise model
- Model is updated as ground test results become available
- Current Best Estimate of LPF is now approaching the NGO requirements!



Requirements noise projection



Current Best Estimate



In Summary

- 🪐 All environmental tests complete
- 🪐 The caging mechanism launch lock drives the critical path
- 🪐 Results from OSTT (TOQM) demonstrate performance better than requirements
- 🪐 Cold Gas thrusters are a viable alternative to the FEEPs.
- 🪐 The project will enter hibernation at the beginning of next year
 - Hibernation is scheduled to last ~15months
- 🪐 Launch is scheduled for June 2014.



Thank you

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ESA ESAC

ESA ESOC

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EADS Astrium GmbH

University of Trento

Albert Einstein Institute

University of Glasgow

University of Birmingham

Imperial College London

ETH Zurich

Institut d'Estudis Espacials de Catalunya

Universidad Politecnica de Barcelona

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